

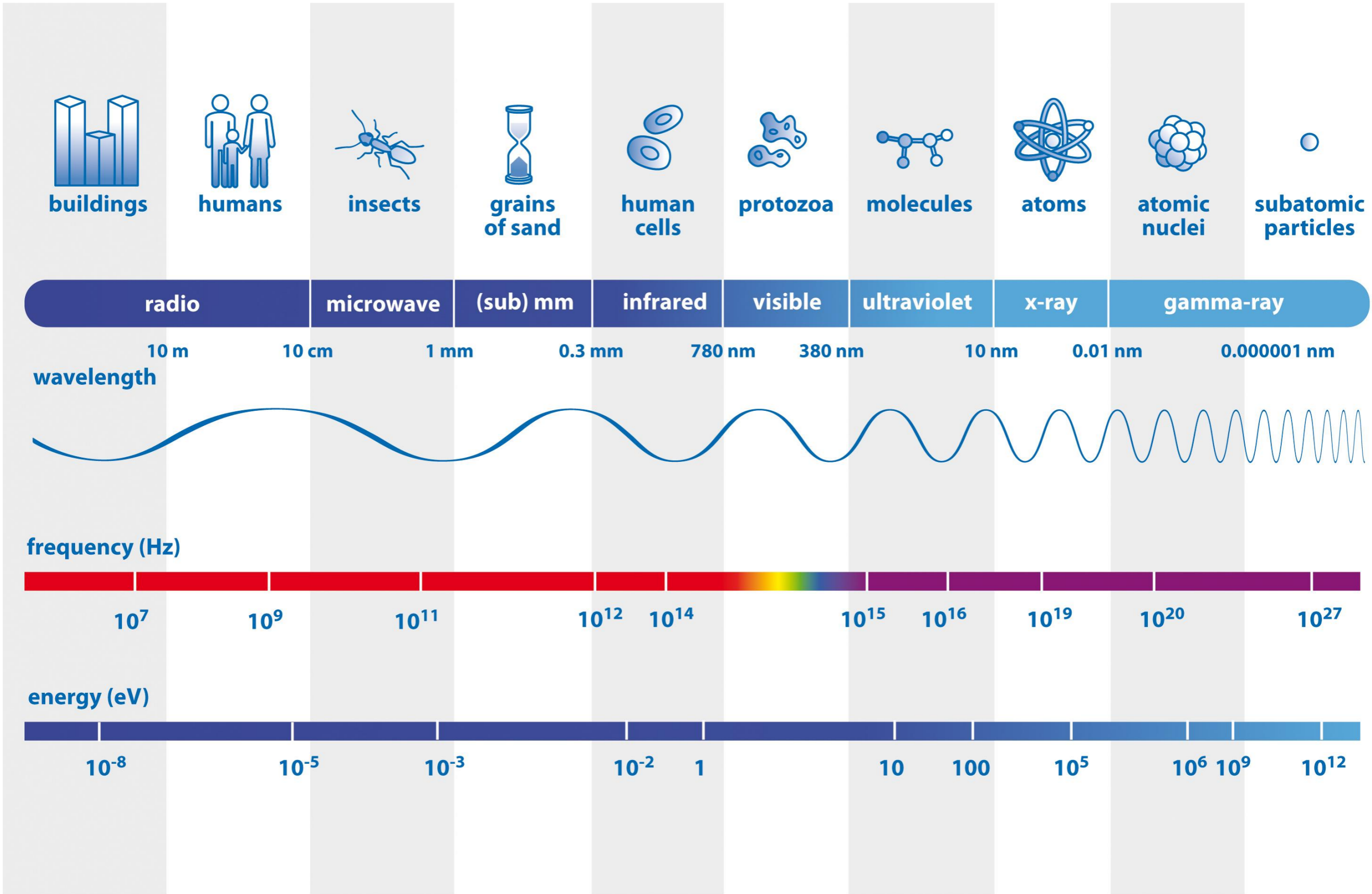
La missione spaziale XMM-Newton per l'astronomia X

Il progetto EXTraS



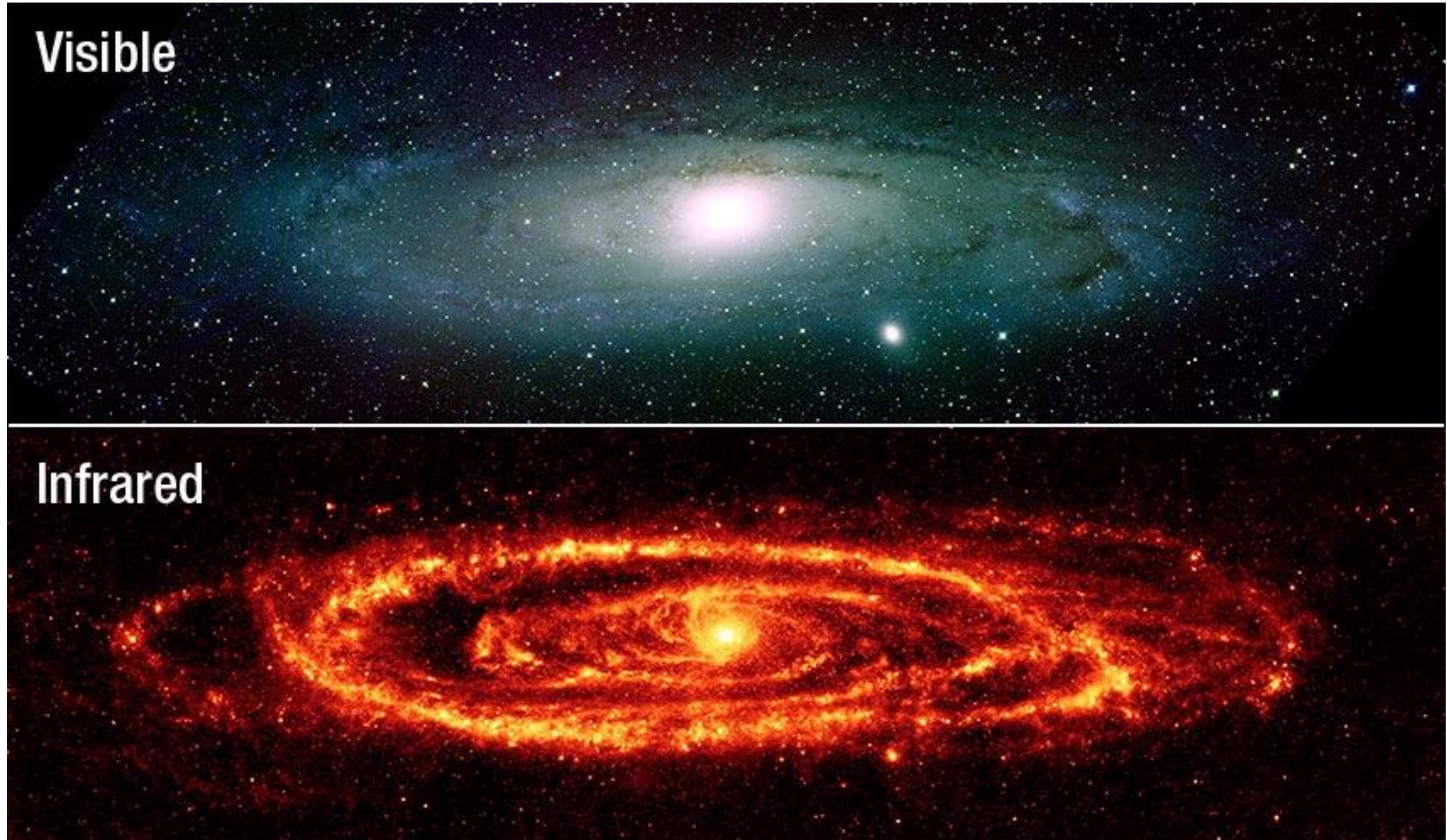
XMM-Newton



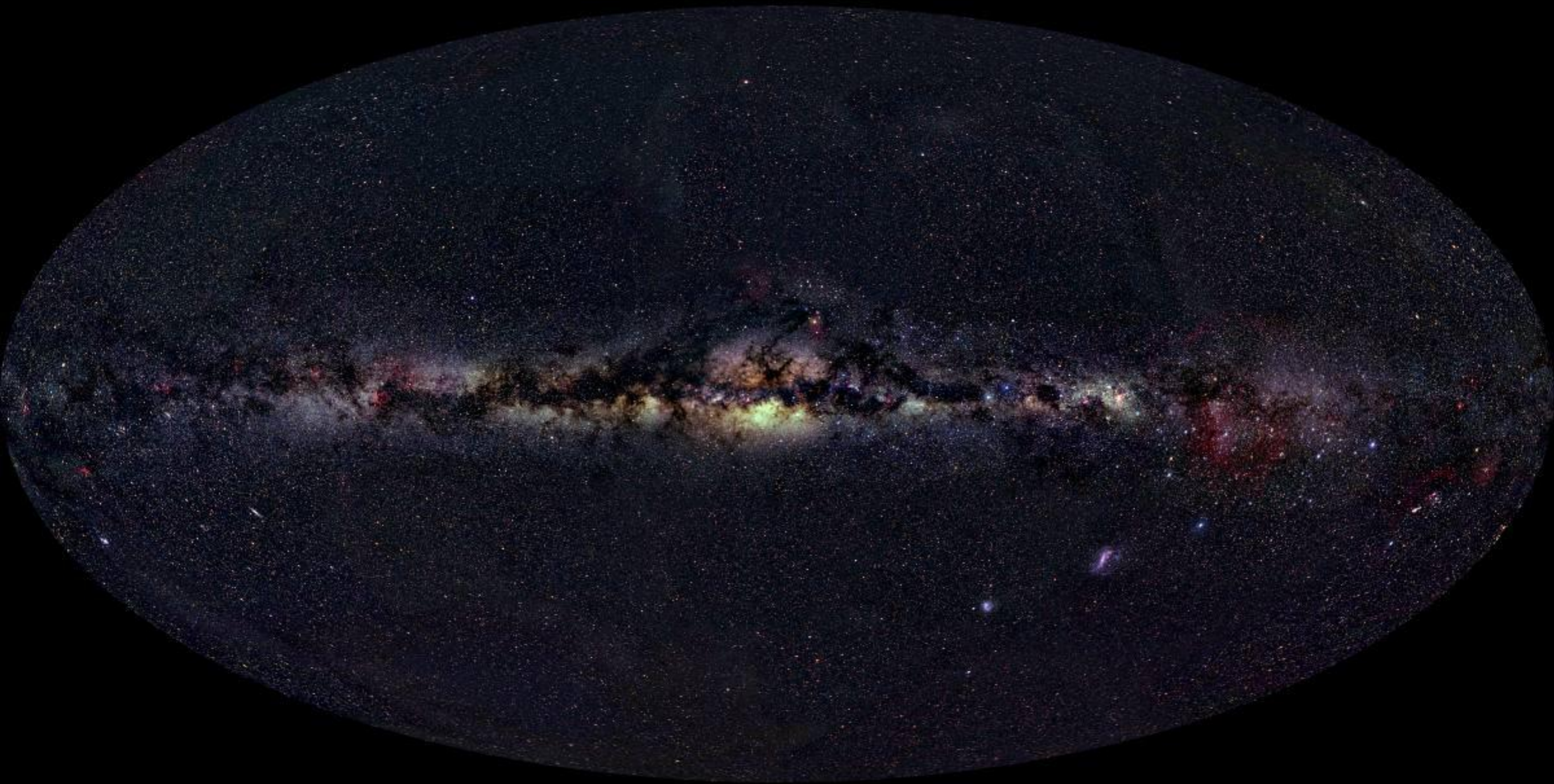




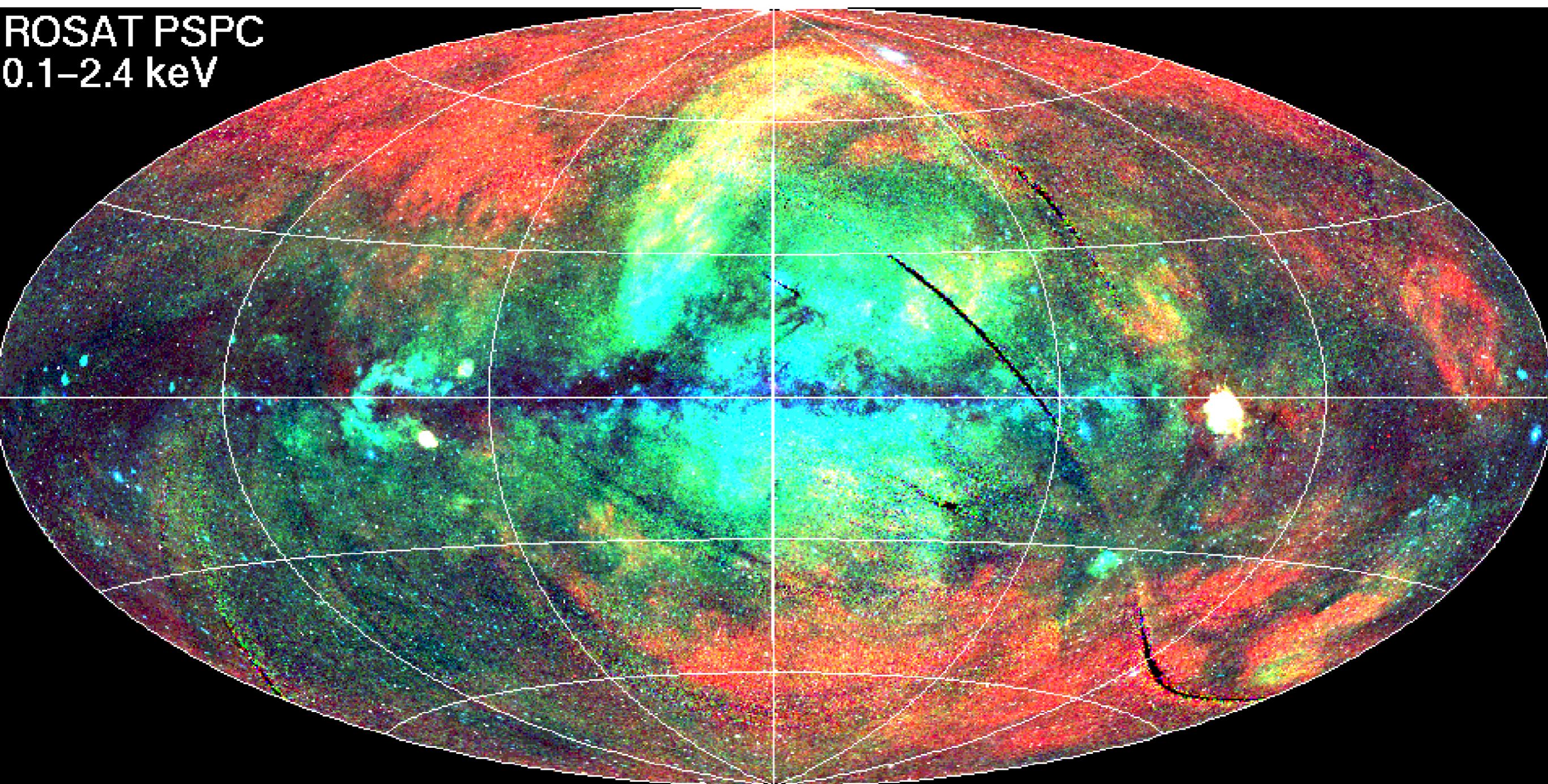
La galassia di Andromeda



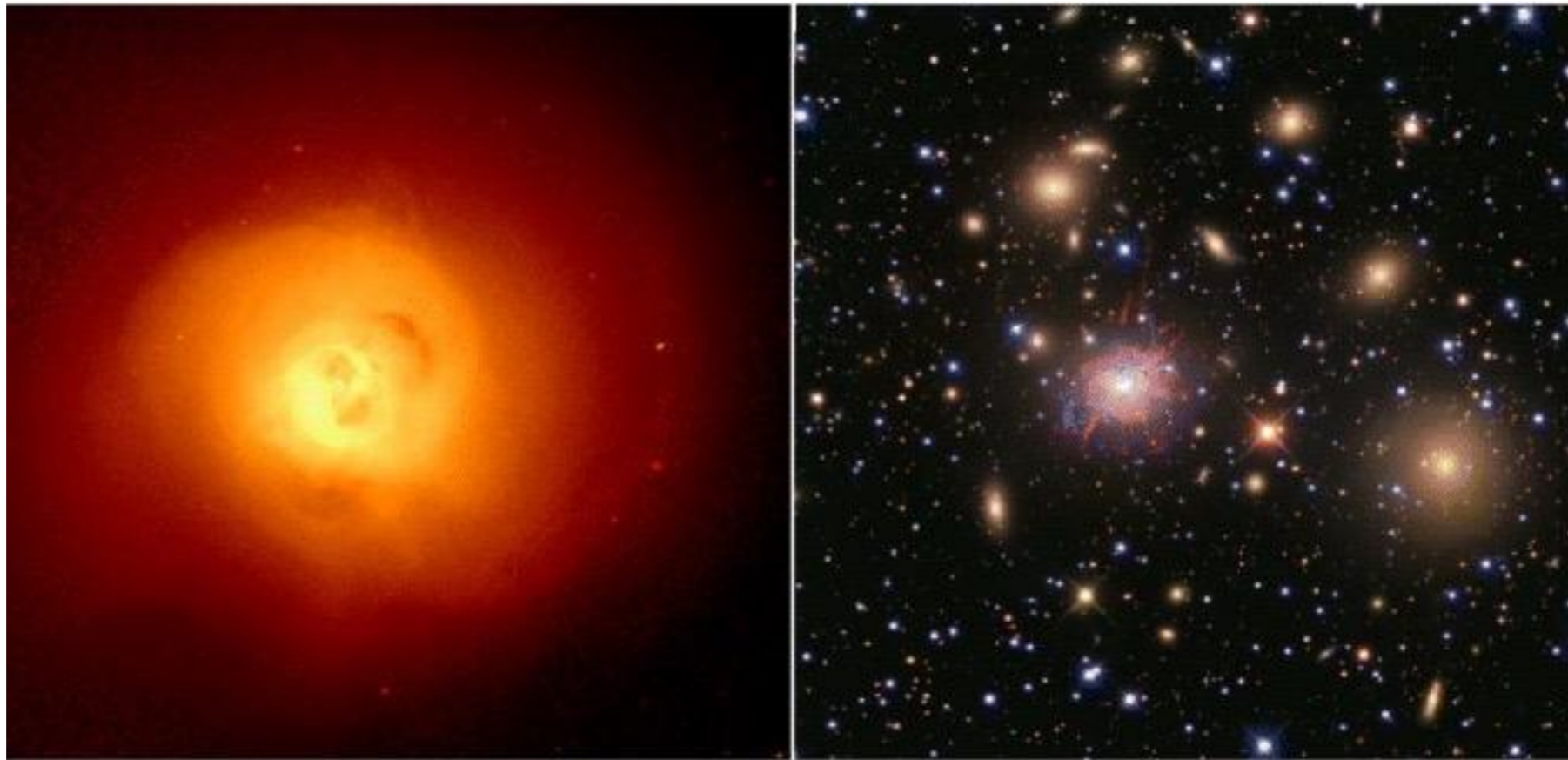
The Deep Sky



ROSAT PSPC
0.1–2.4 keV

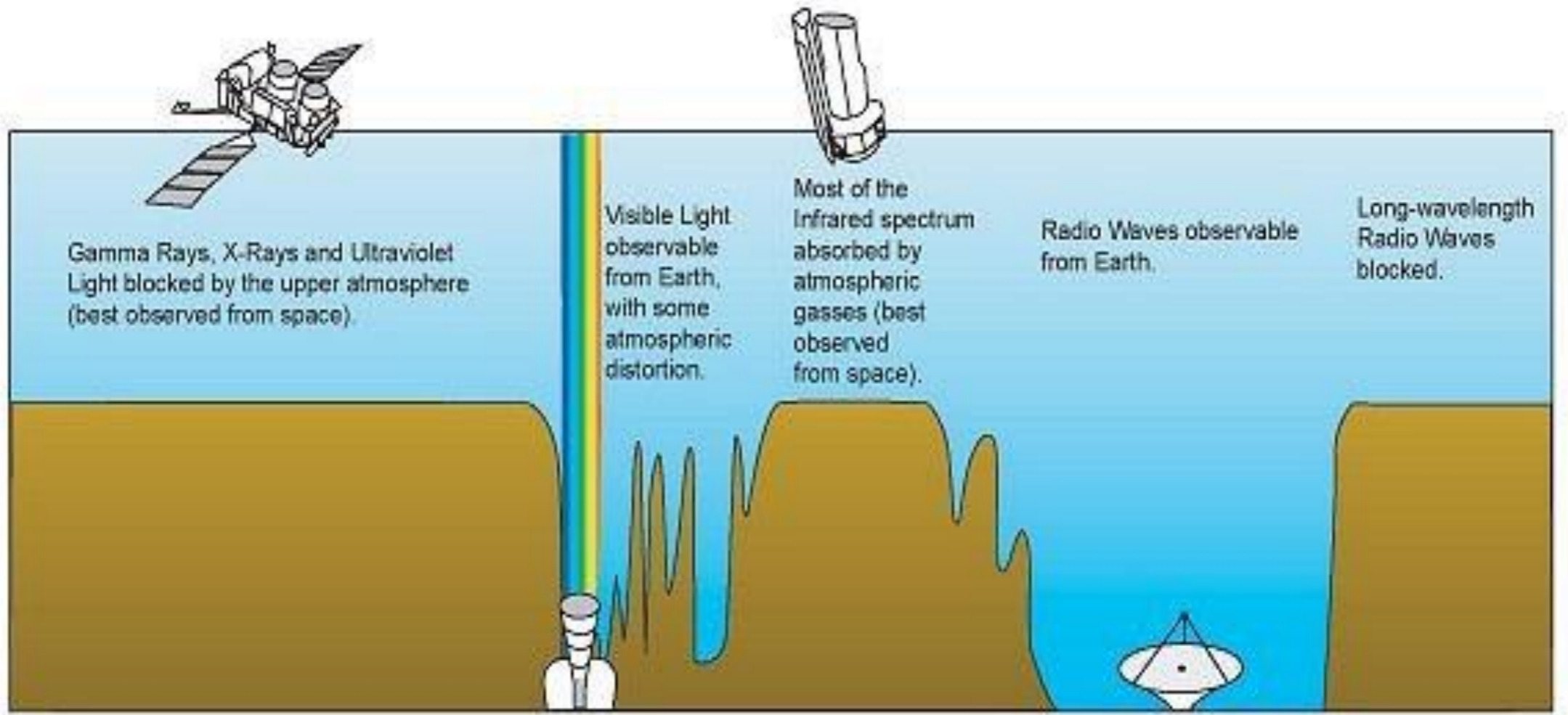
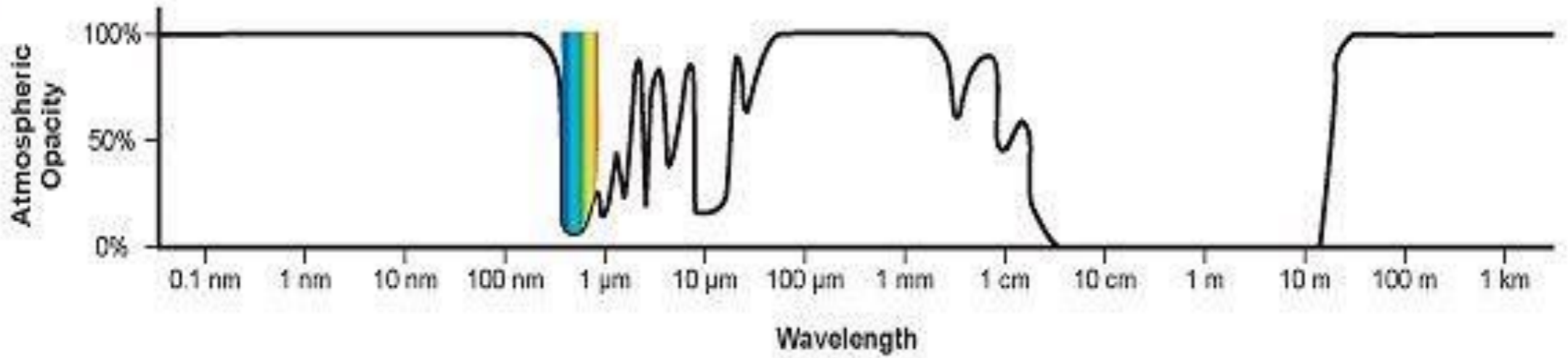


Perseus cluster of galaxies



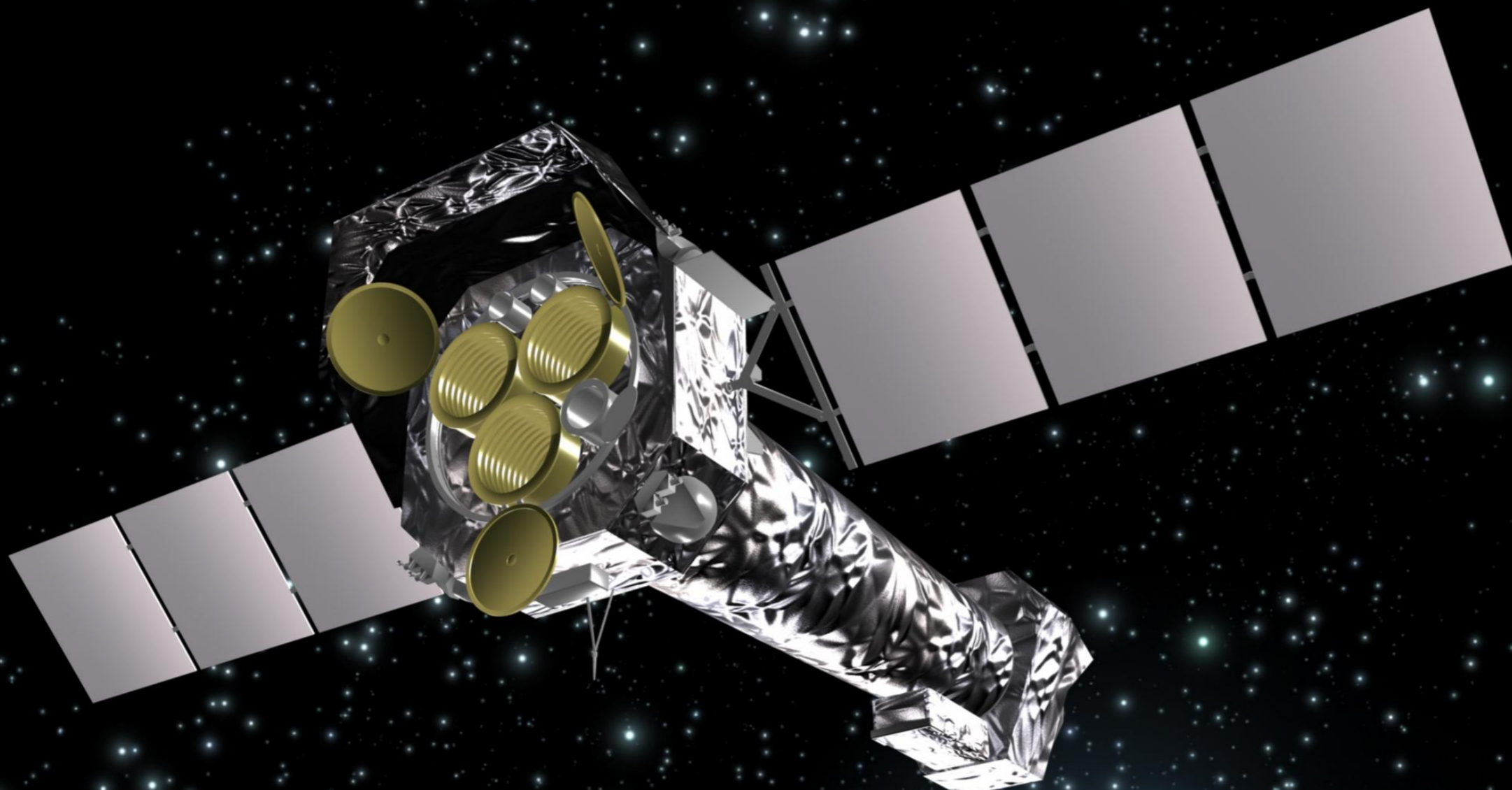
X-rays

visible light



La missione XMM-Newton

- proposed in 1984
- approved in 1985;
- project team was formed in 1993
- development work began in 1996.
- constructed and tested from March 1997 to September 1999.
- launched in Dec 1999
- in-orbit commissioning started Jan 2000
- first images were published Feb 2000
- still fully operational



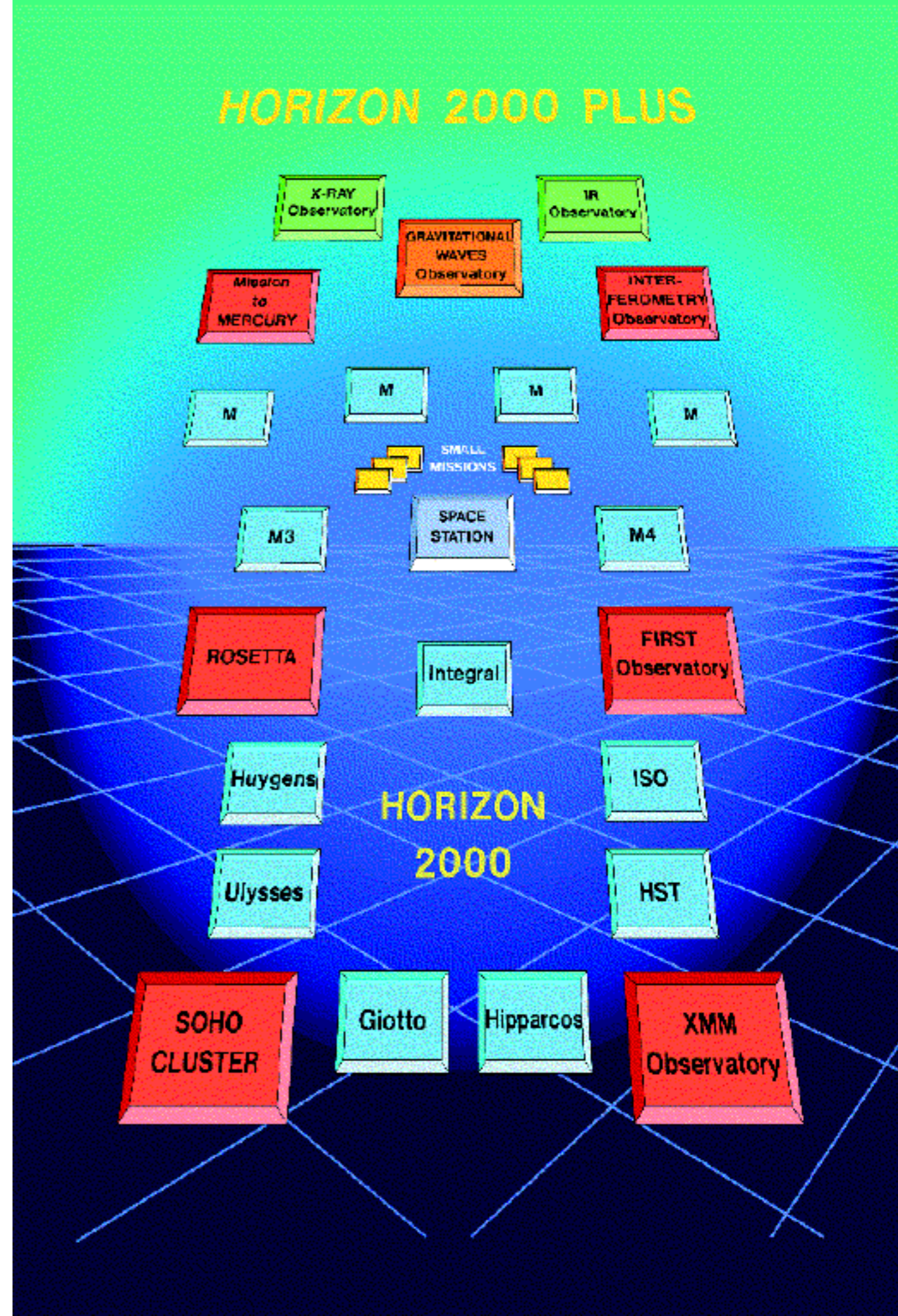
Cornerstone mission

nel programma di lungo termine

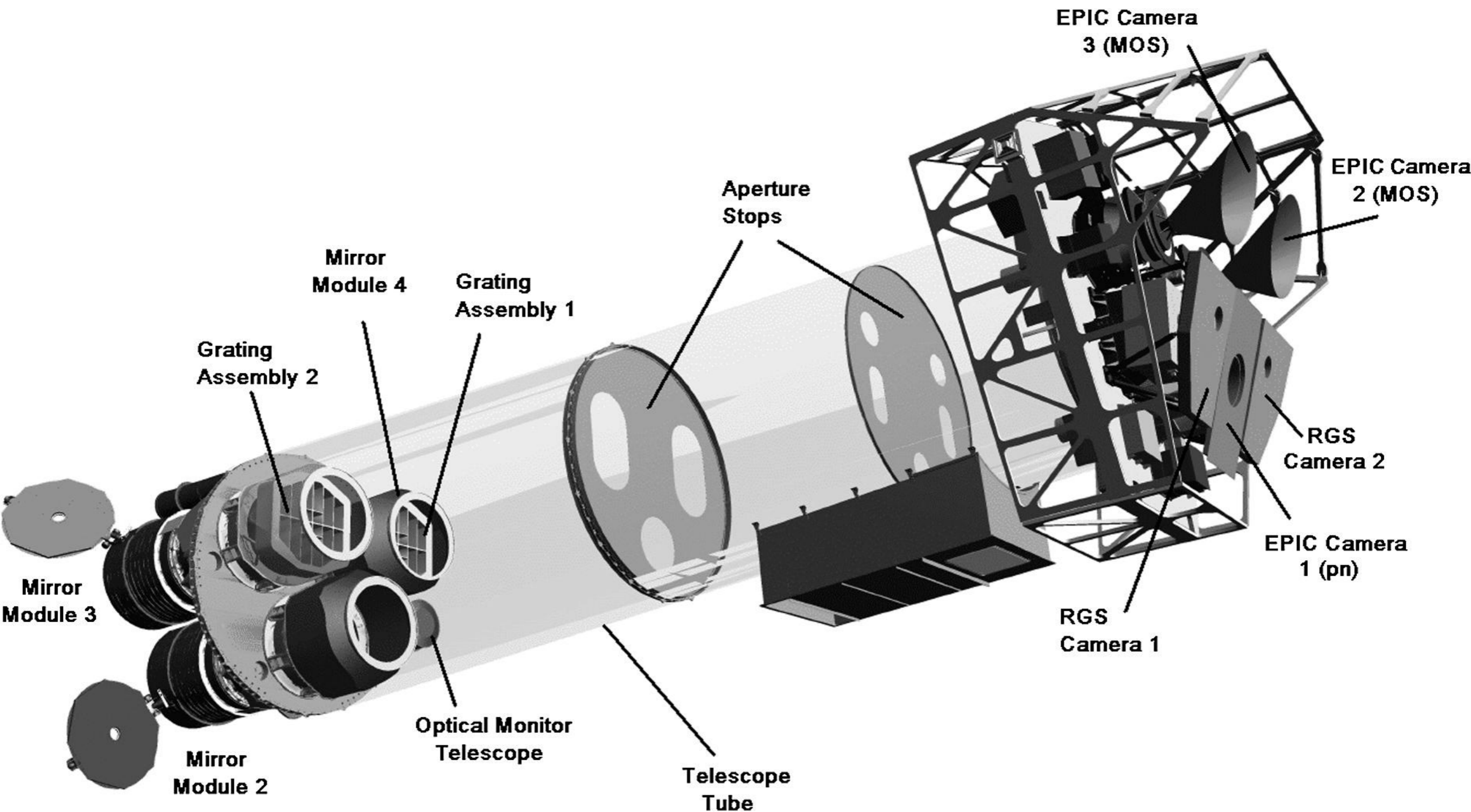
Horizon 2000 dell'ESA

<http://xmm.esac.esa.int>

<http://sci.esa.int/xmm-newton/>

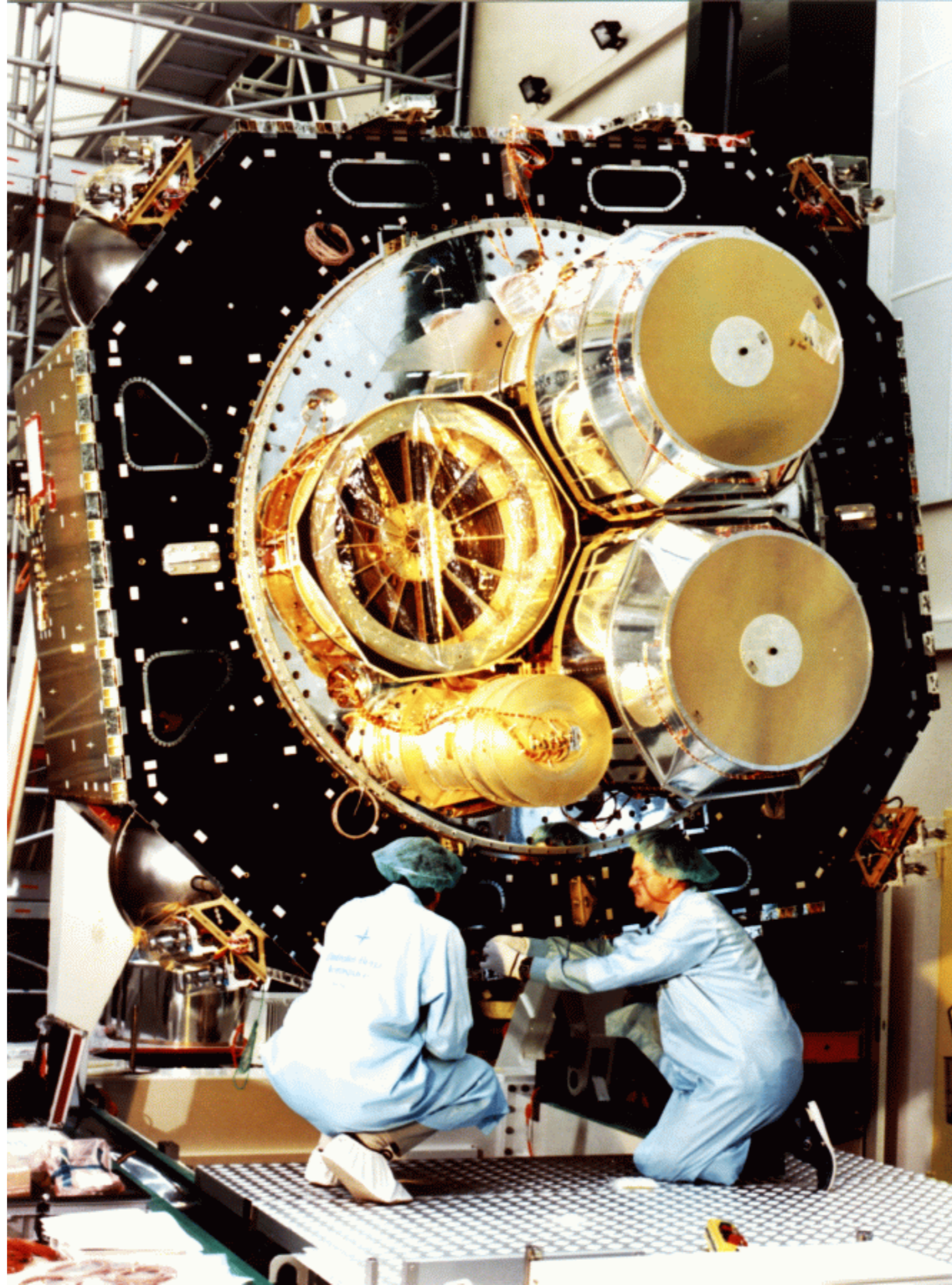


Come e' fatto XMM-Newton



XMM 3D model

<http://sci.esa.int/xmm-newton/31382-3d-model/>

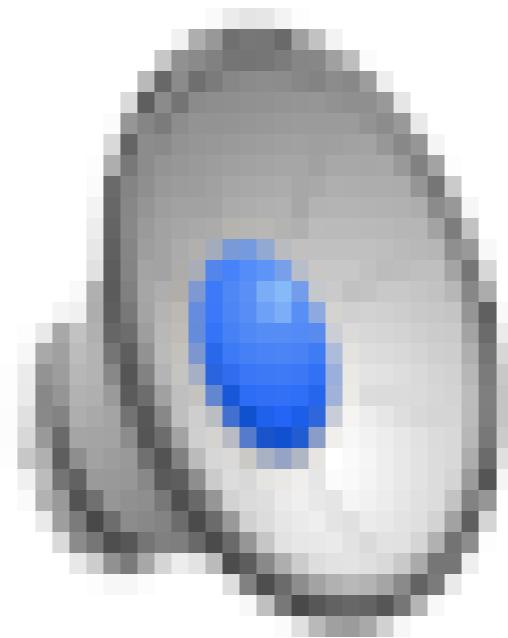


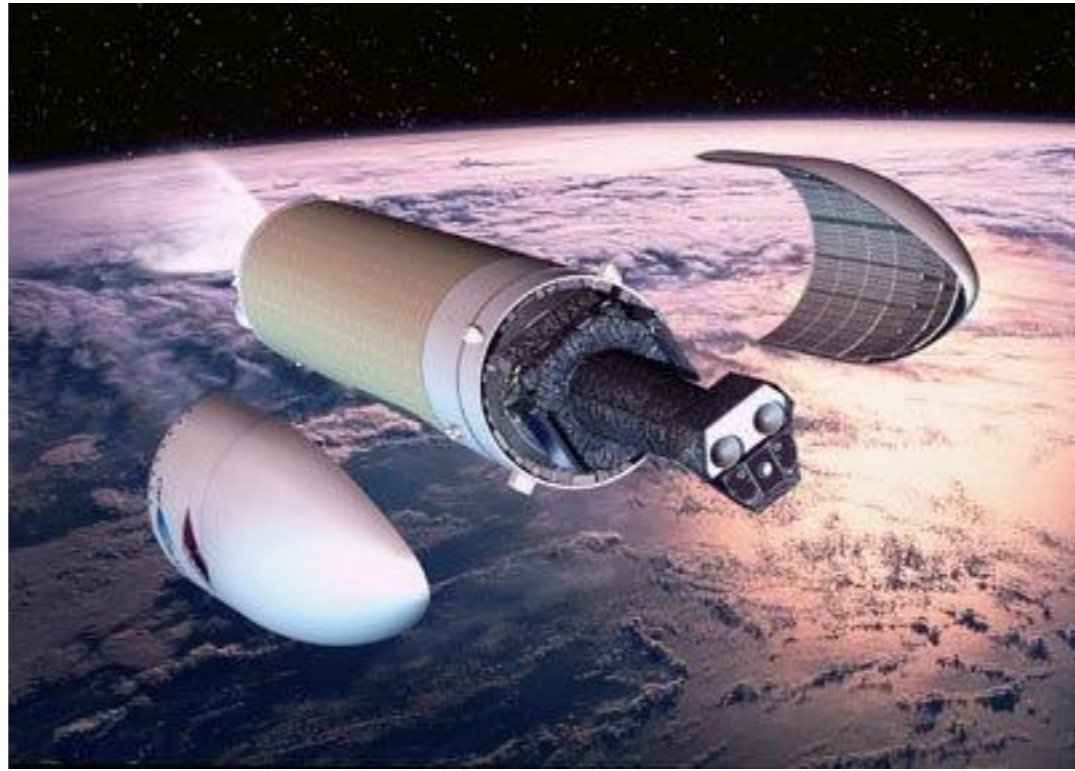


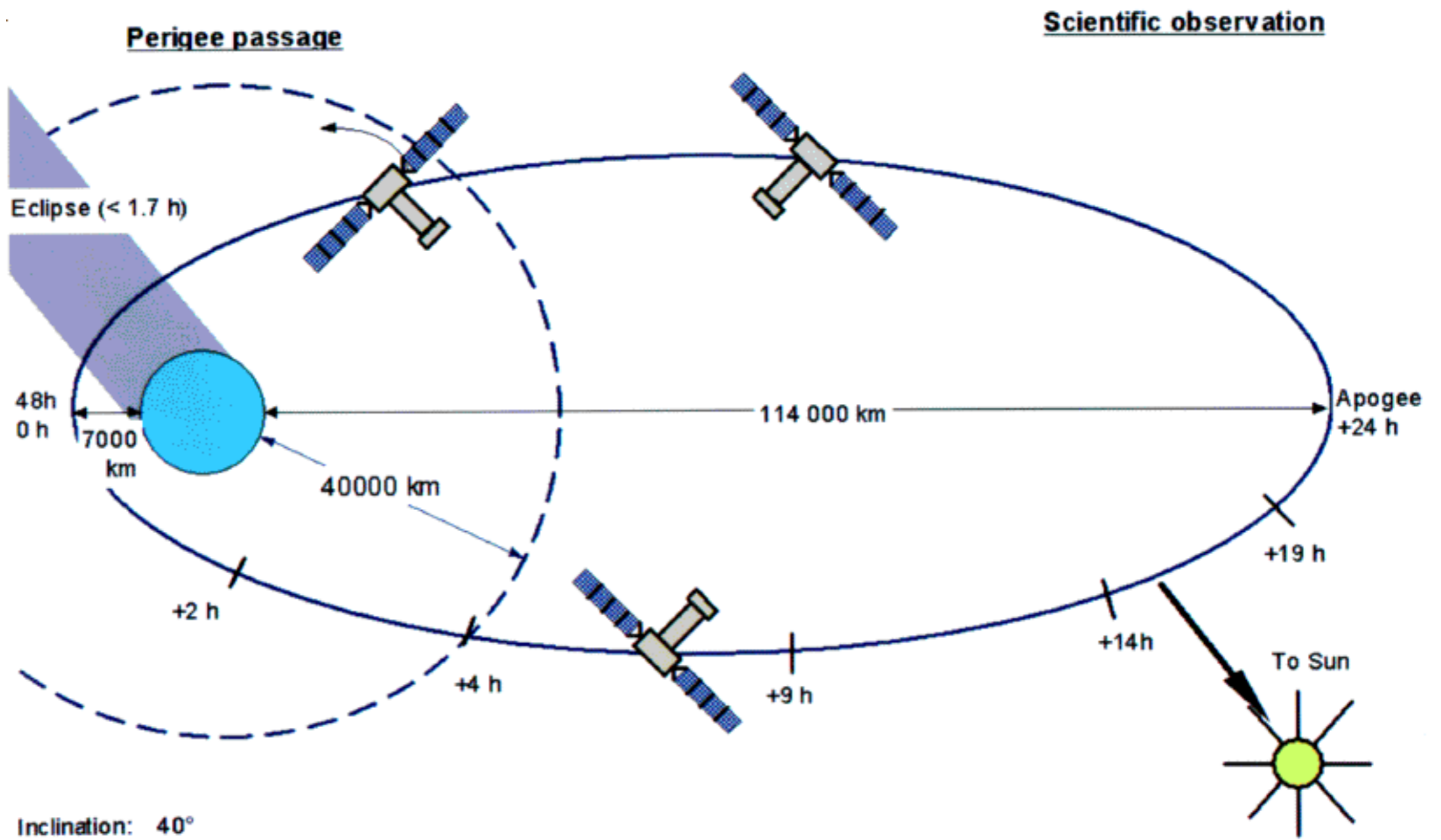


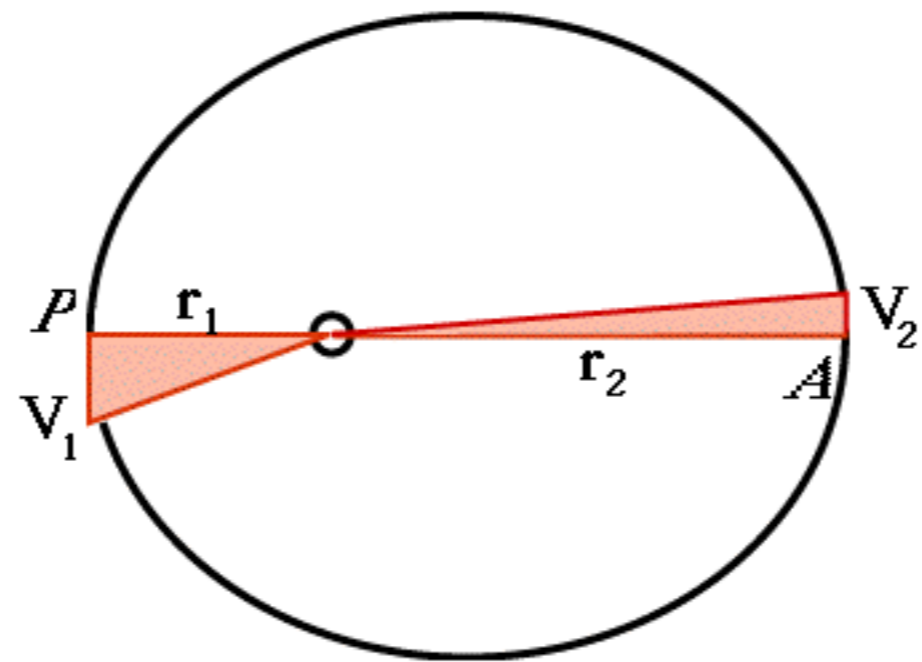
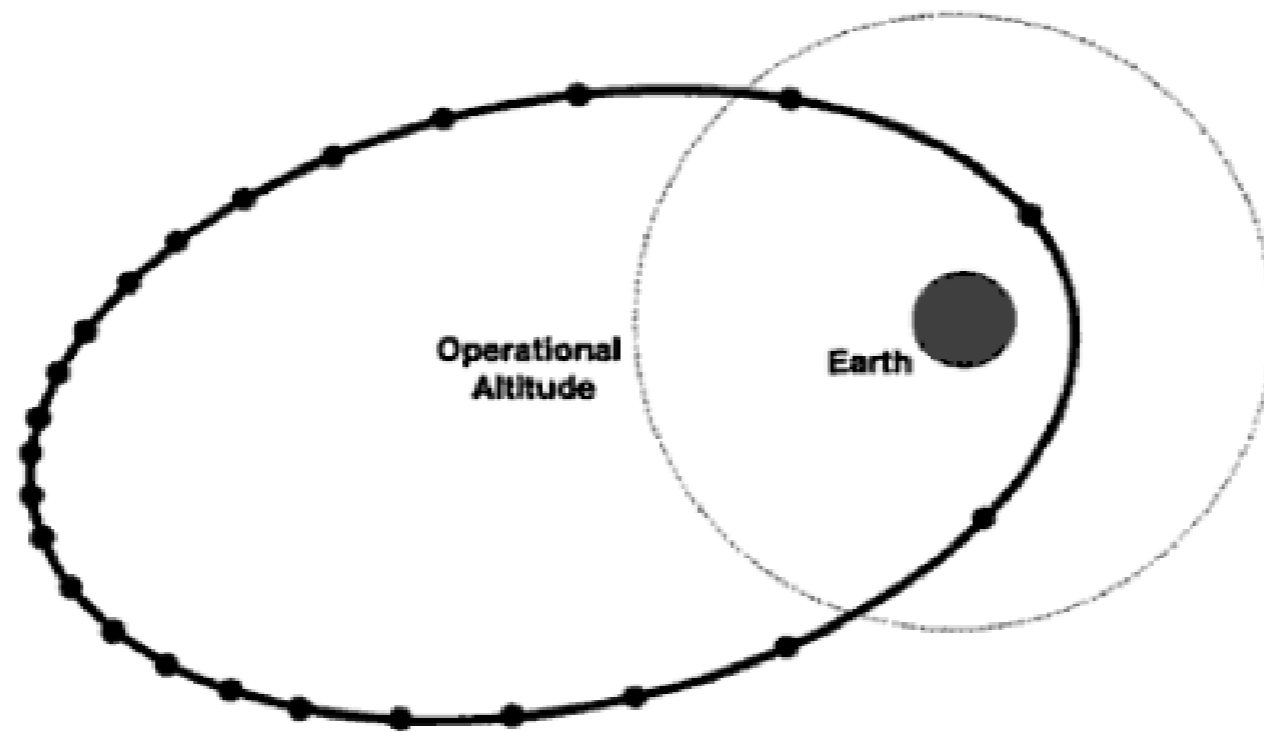
Video time: the launch

http://www.esa.int/spaceinvideos/Videos/2005/05/Stories_from_XMM-English

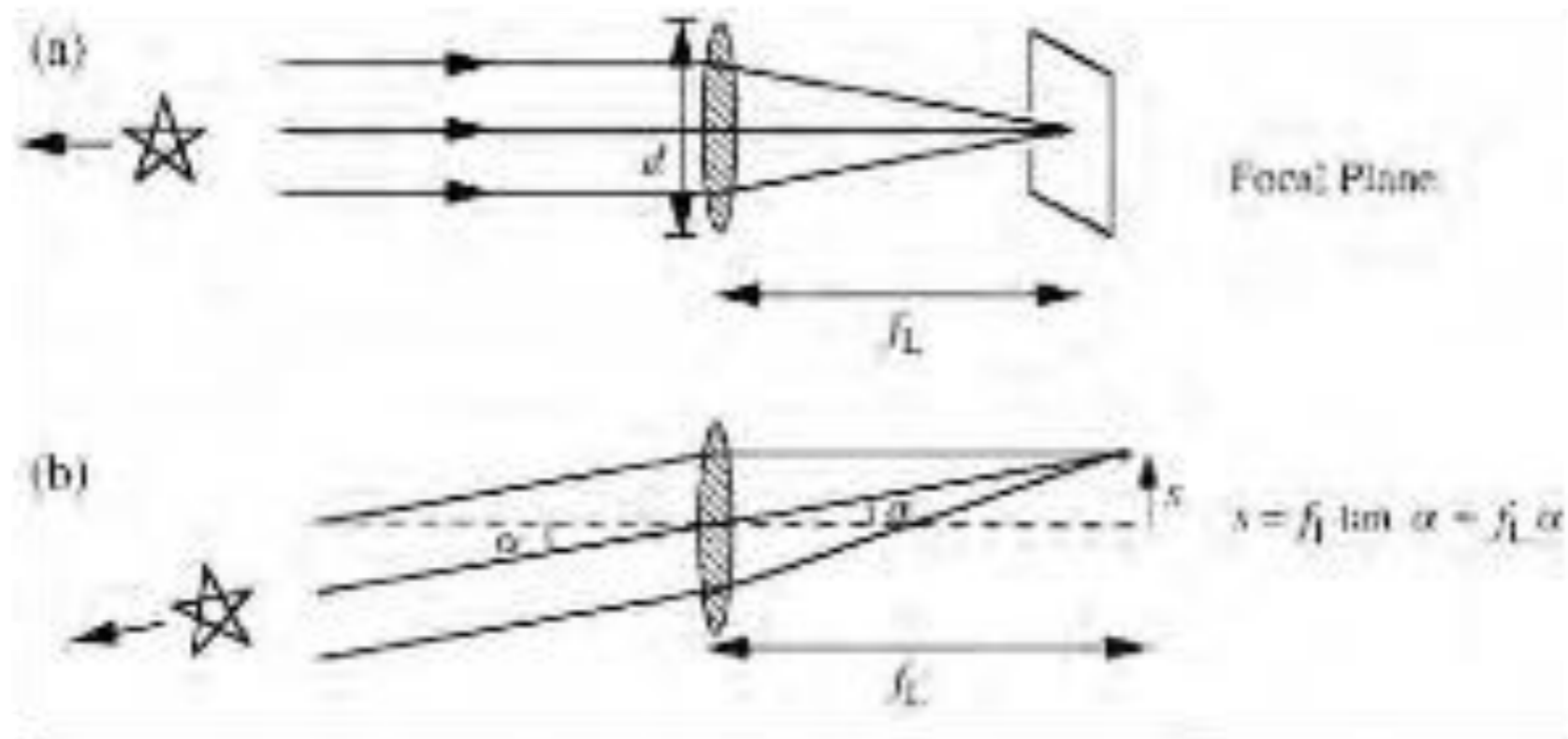




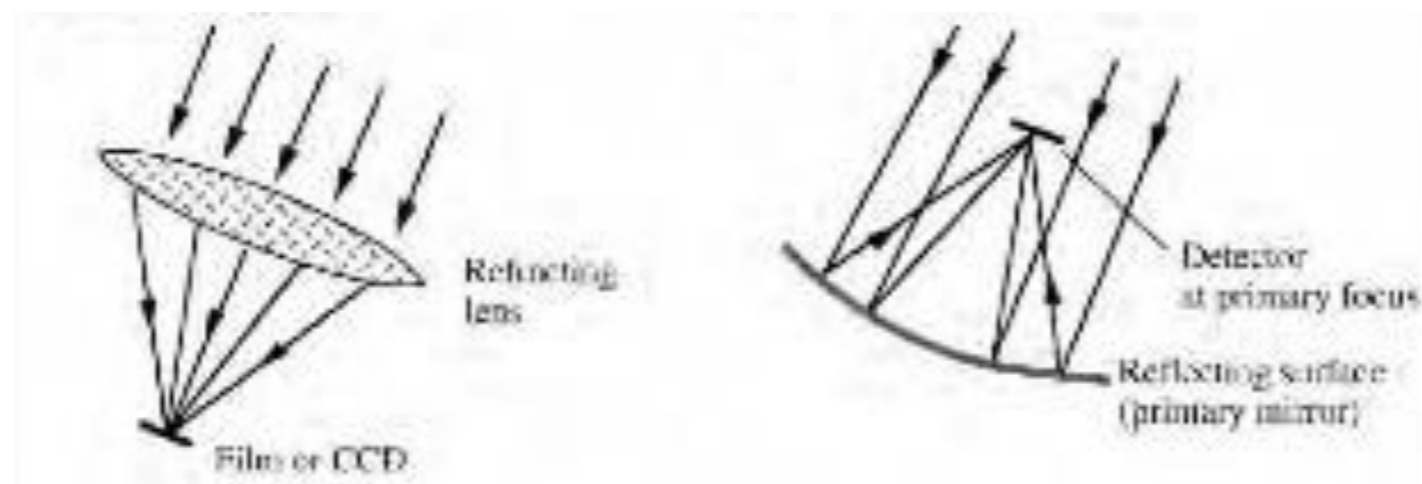




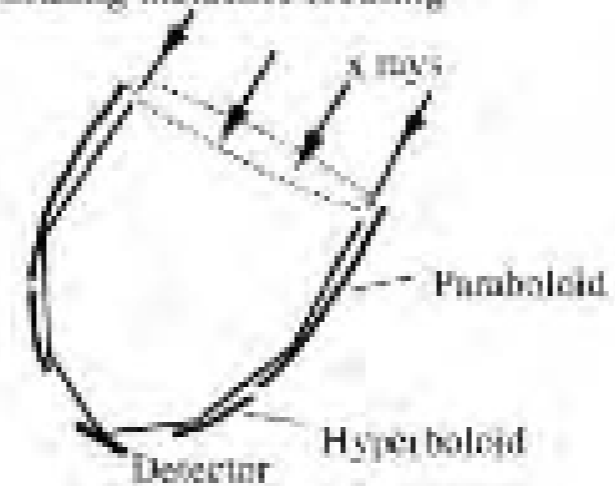
Formazione immagini in un telescopio



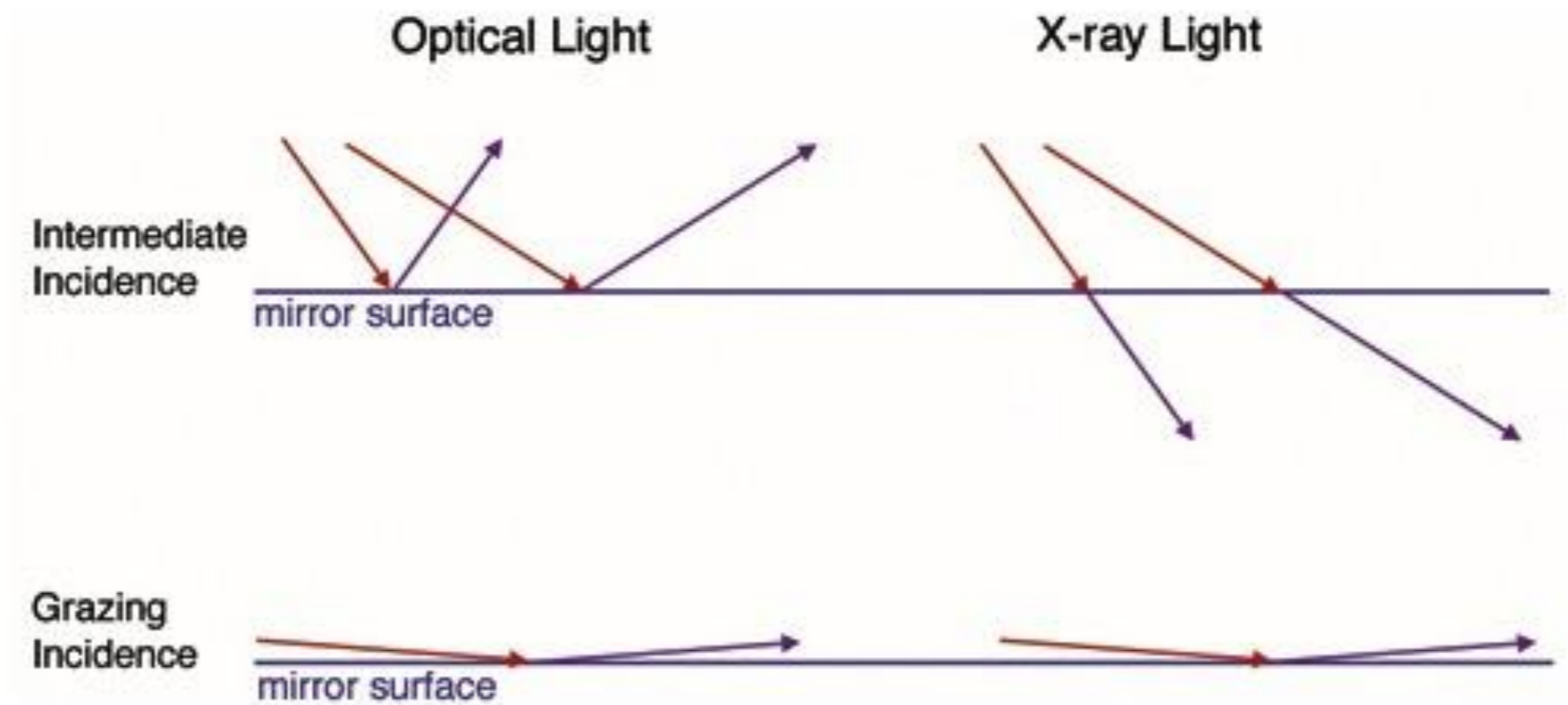
Diversi disegni ottici per diversi telescopi



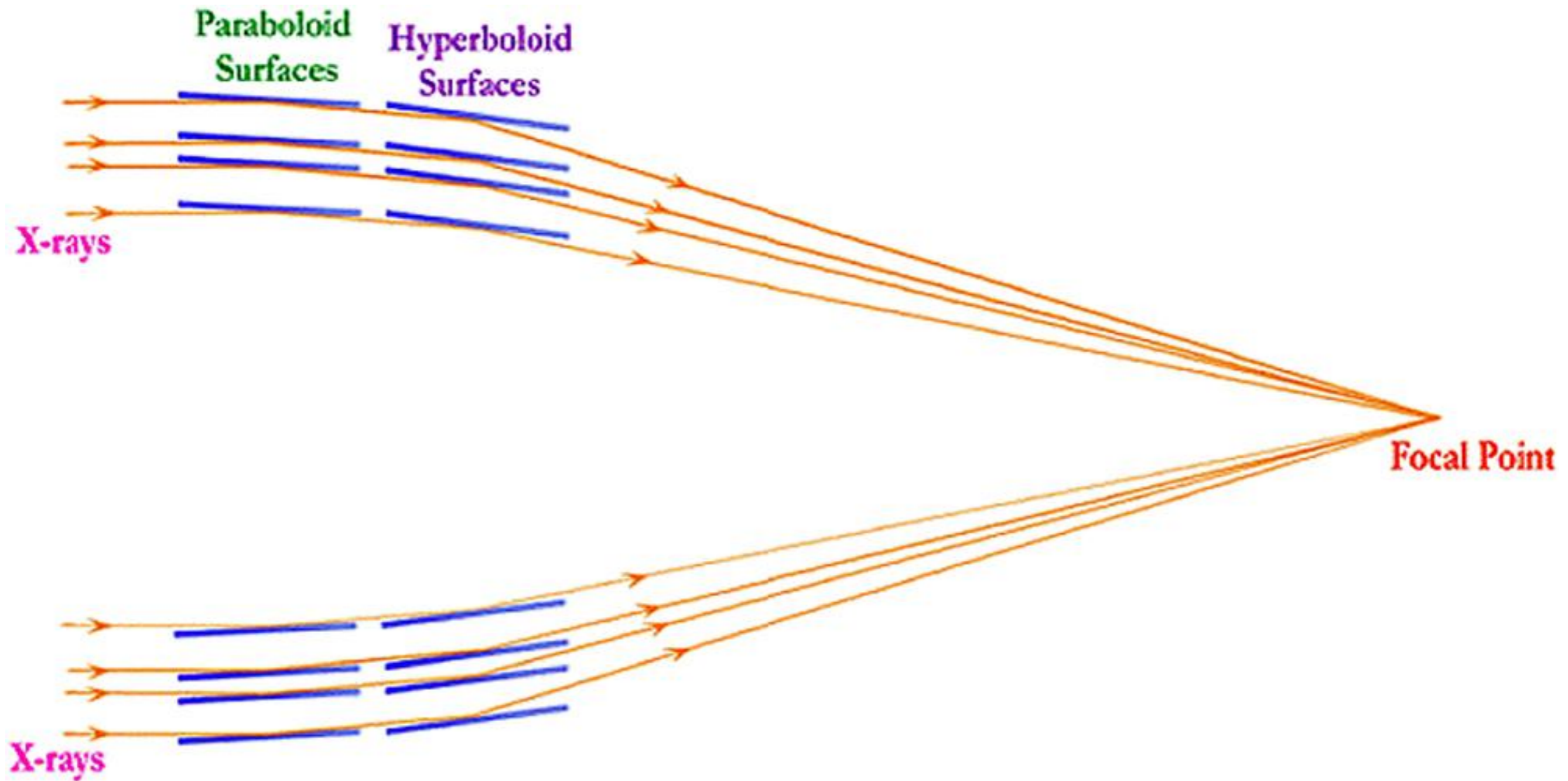
(f) Grazing-incidence focusing



Come focalizzare la luce X



Il tipico disegno di un telescopio X



Video time! X-ray vision


http://www.esa.int/spaceinvideos/Videos/2005/05/Stories_from_XMM-English

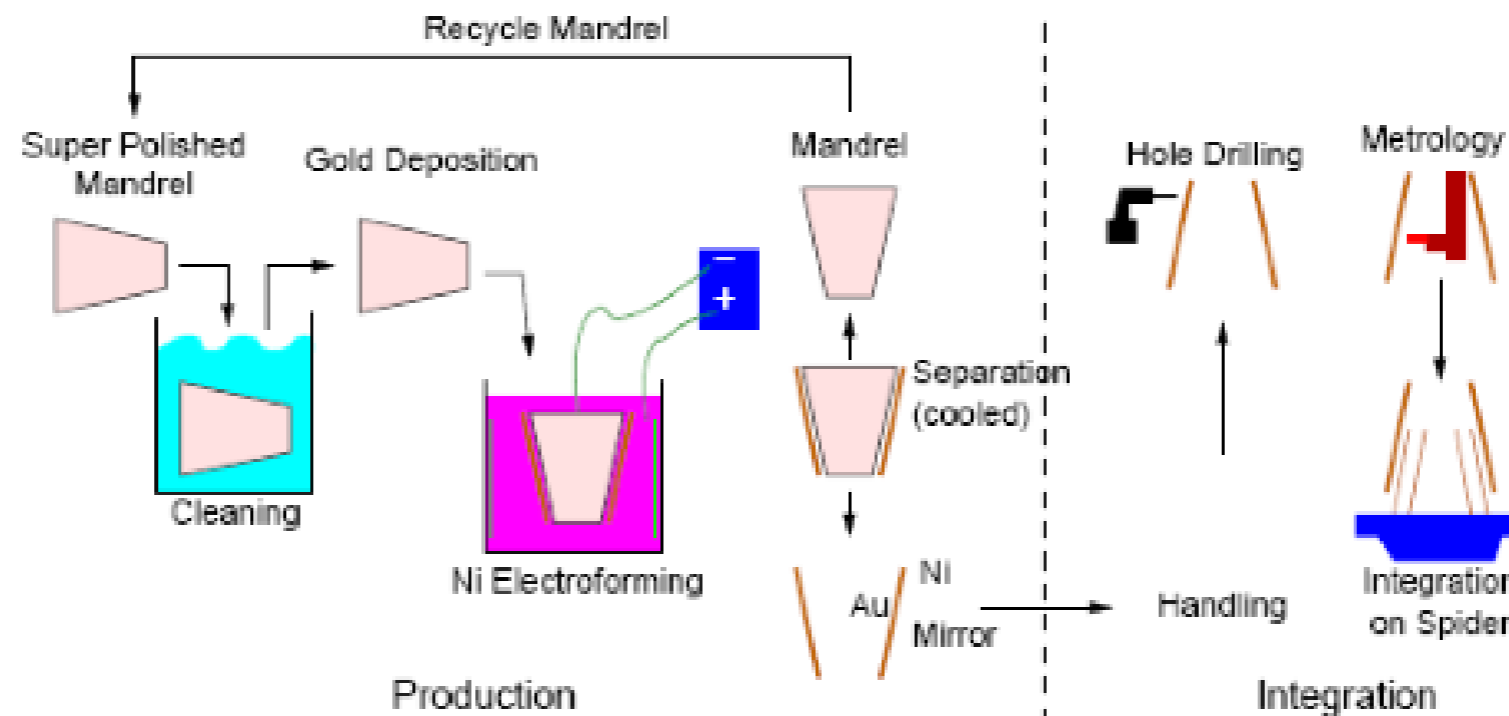




XMM-Newton mirrors during integration

Image courtesy of Doornik Satellitensysteme GmbH

European Space Agency 



(after ESA)

Recipe for making an X-ray mirror:

1. Produce mirror negative ("Mandrels"): Al coated with Kanigen nickel (Ni+10% phosphorus), super-polished [0.4 nm roughness]).
2. Deposit 250 nm Au onto Mandrel
3. Deposit 1 mm Ni onto mandrel ("electro-forming", 10 $\mu\text{m}/\text{h}$)
4. Cool Mandrel with liquid N. Au sticks to Nickel
5. Verify mirror on optical bench.

Total production time of one mirror: 12 d, for XMM: 3 \times 58 mirrors.

Video time!

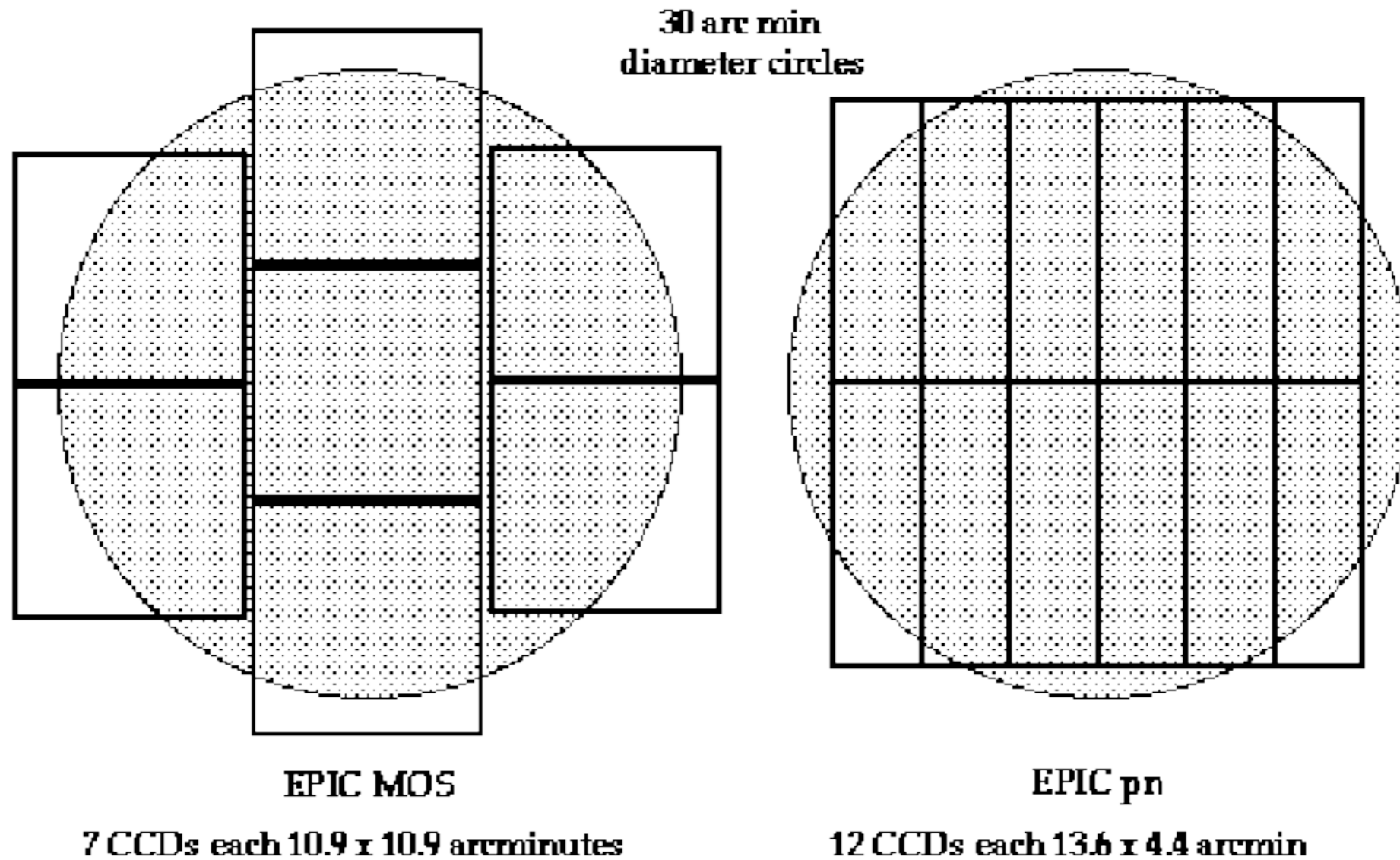
Golden eyes

http://www.esa.int/spaceinvideos/Videos/2005/05/Stories_from_XMM-English

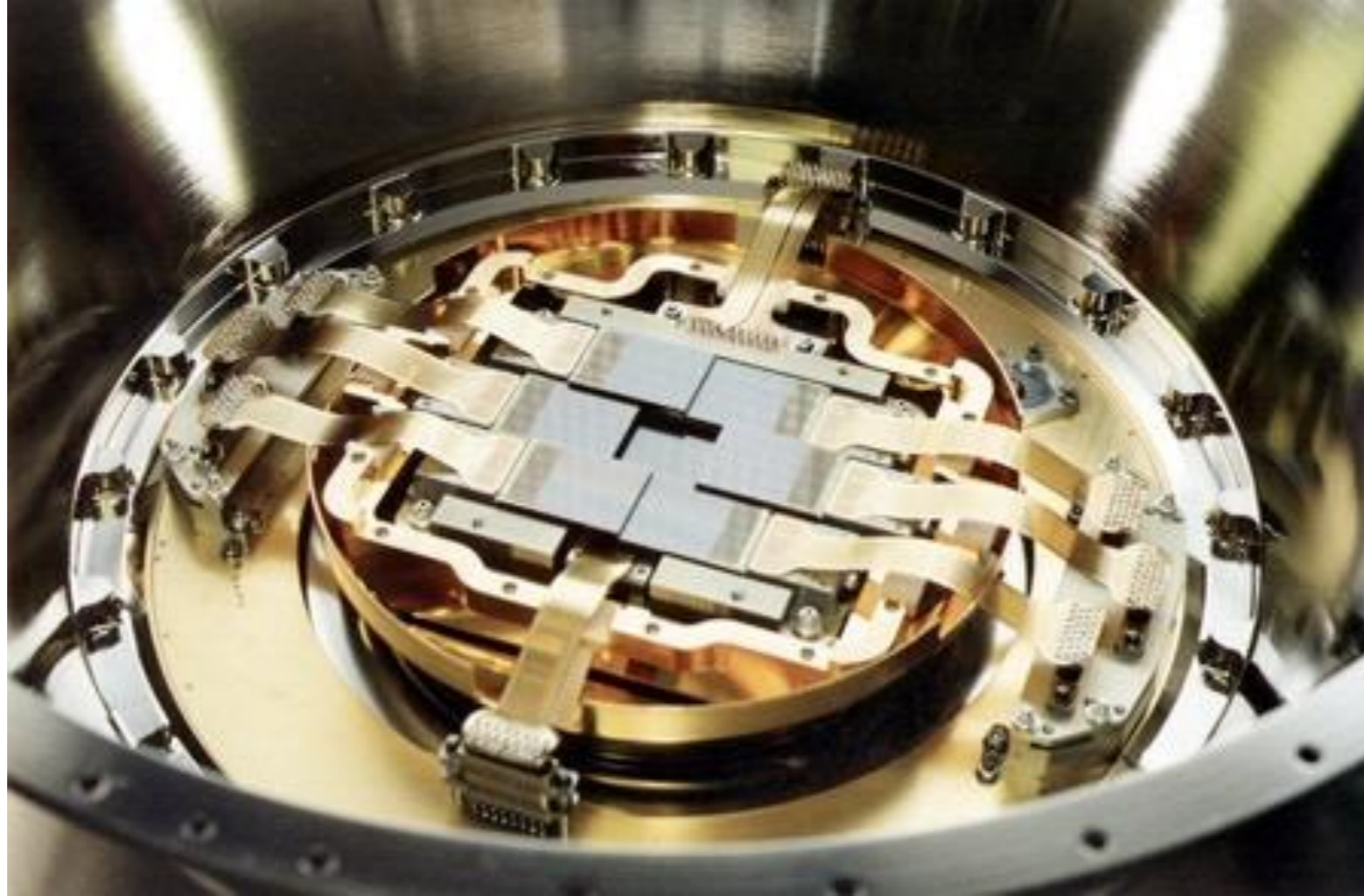


Lo strumento EPIC di XMM-Newton

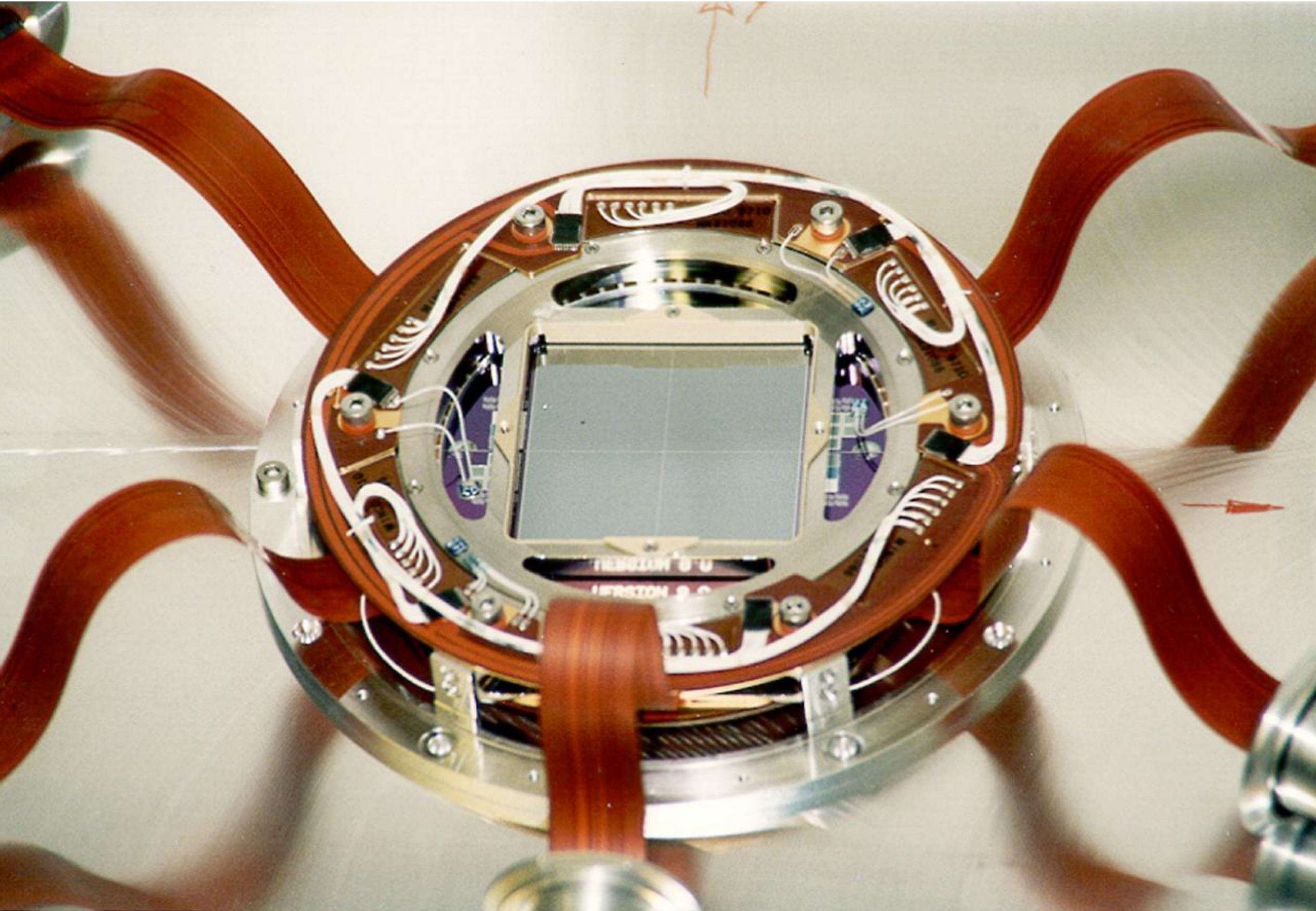
Comparison of focal plane organisation of EPIC MOS and pn cameras



Una delle camere MOS



La camera pn



Come funziona
un CCD
per astronomia X

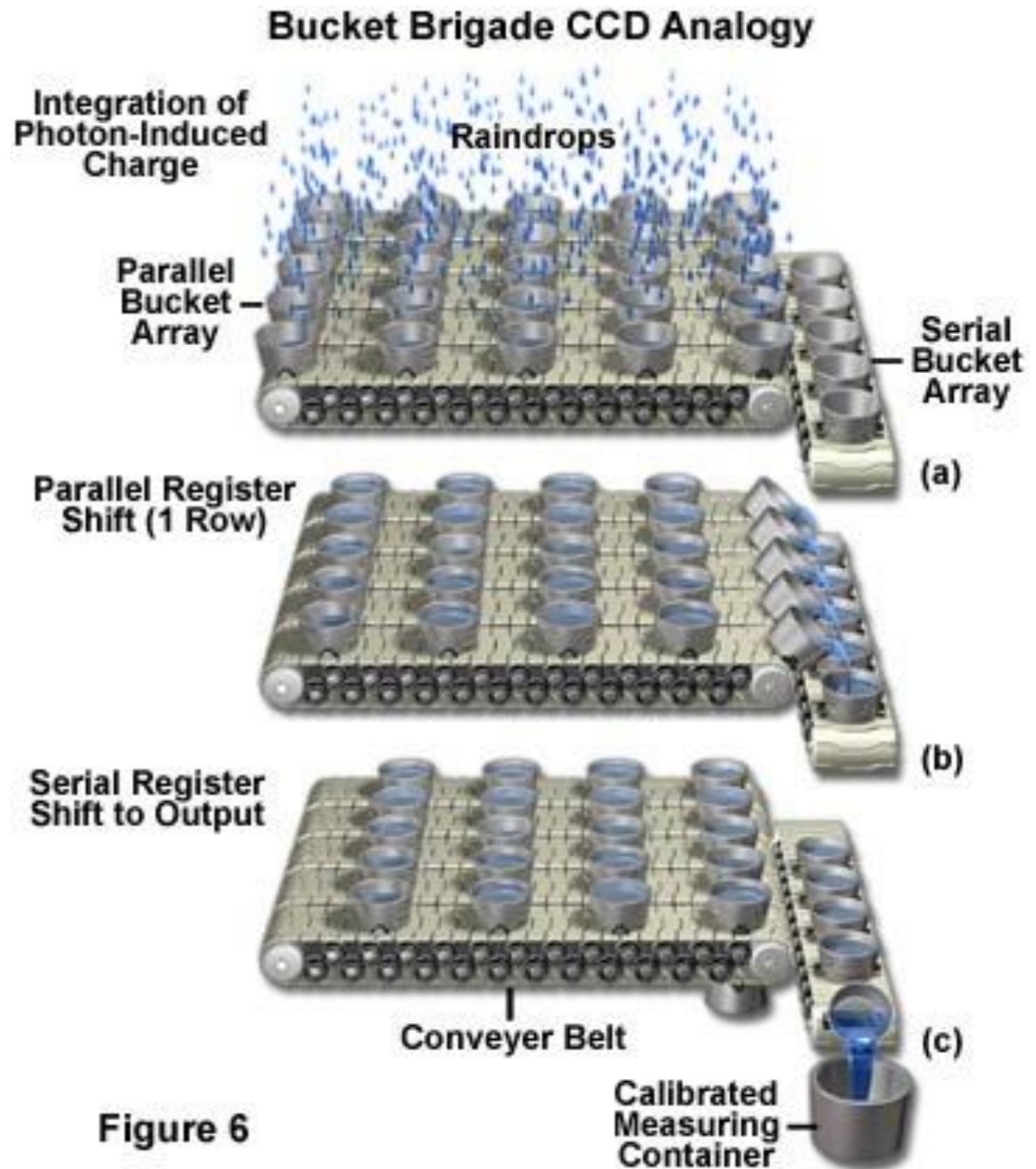
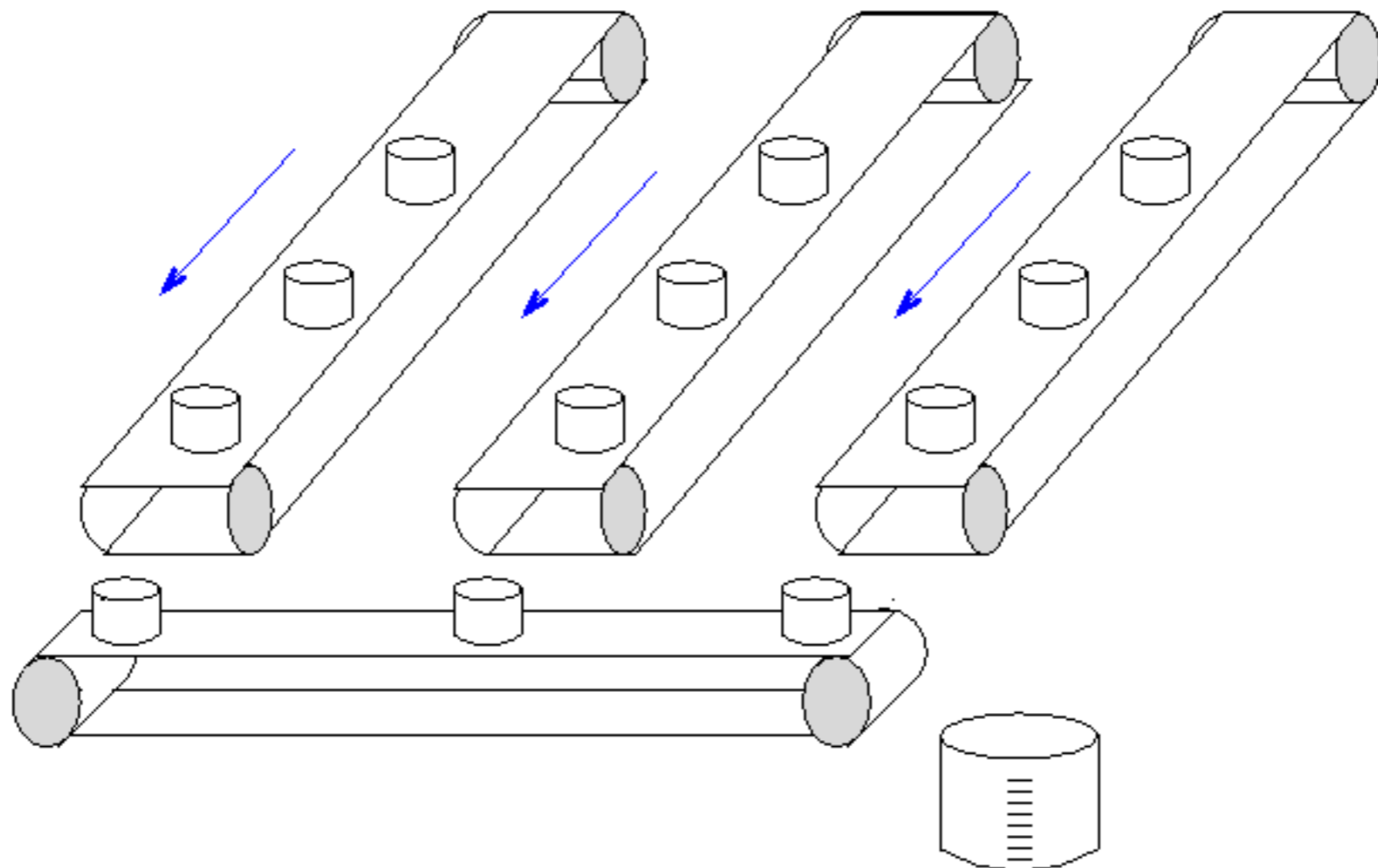


Figure 6



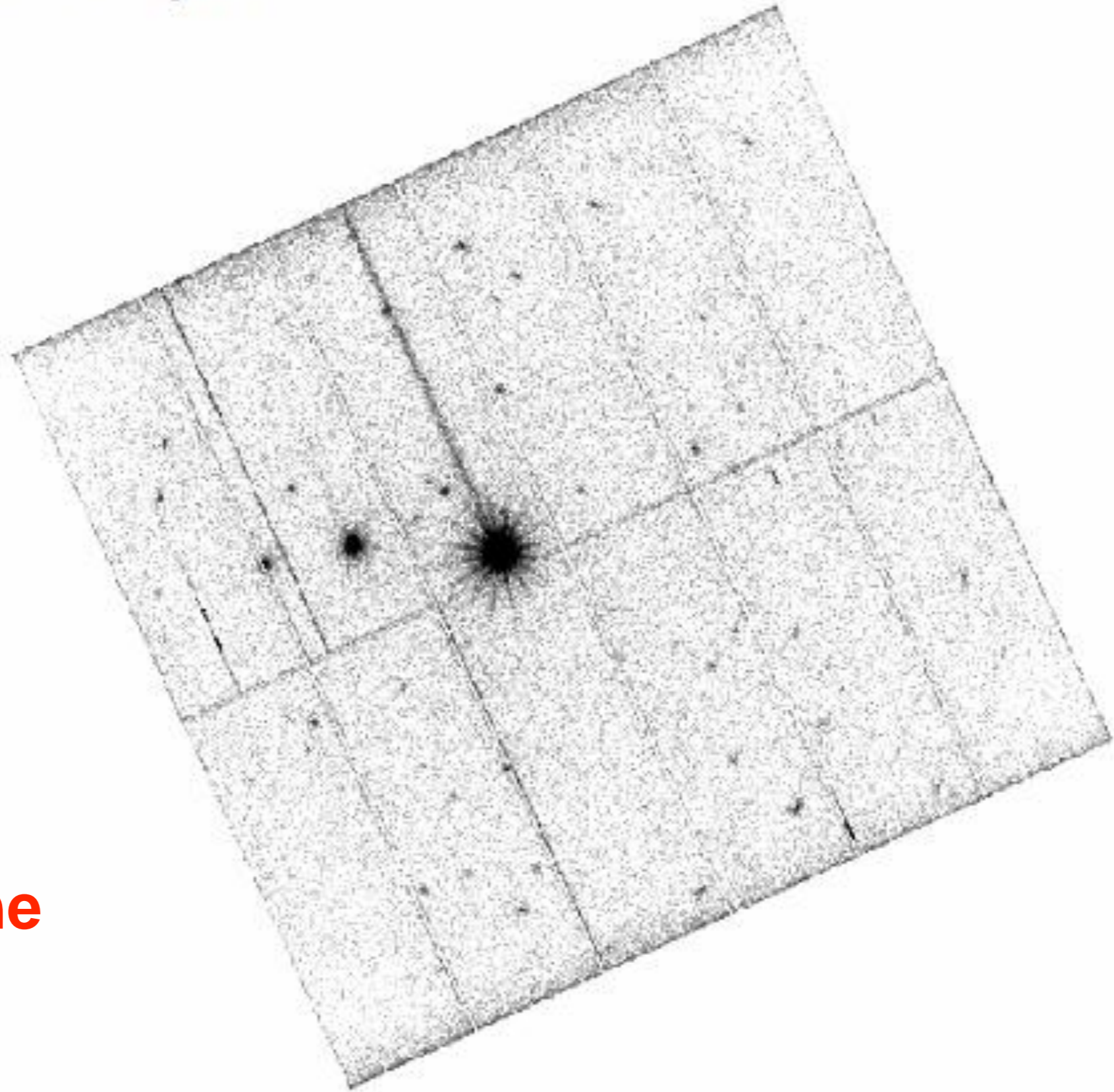
- Strumento con capacità imaging (CCD)
- Raccolgo lista di “eventi”
- Un evento: t, E, x, y

- Eventi :
 1. fotoni, da sorgenti (puntiformi, diffuse, fondo diffuso)
 2. non-fotoni (background “strumentale”)

- Energia dei fotoni
- Tempo di arrivo
- Posizione (direzione di arrivo)

Come sfruttiamo queste informazioni?

XMM EPIC/pn

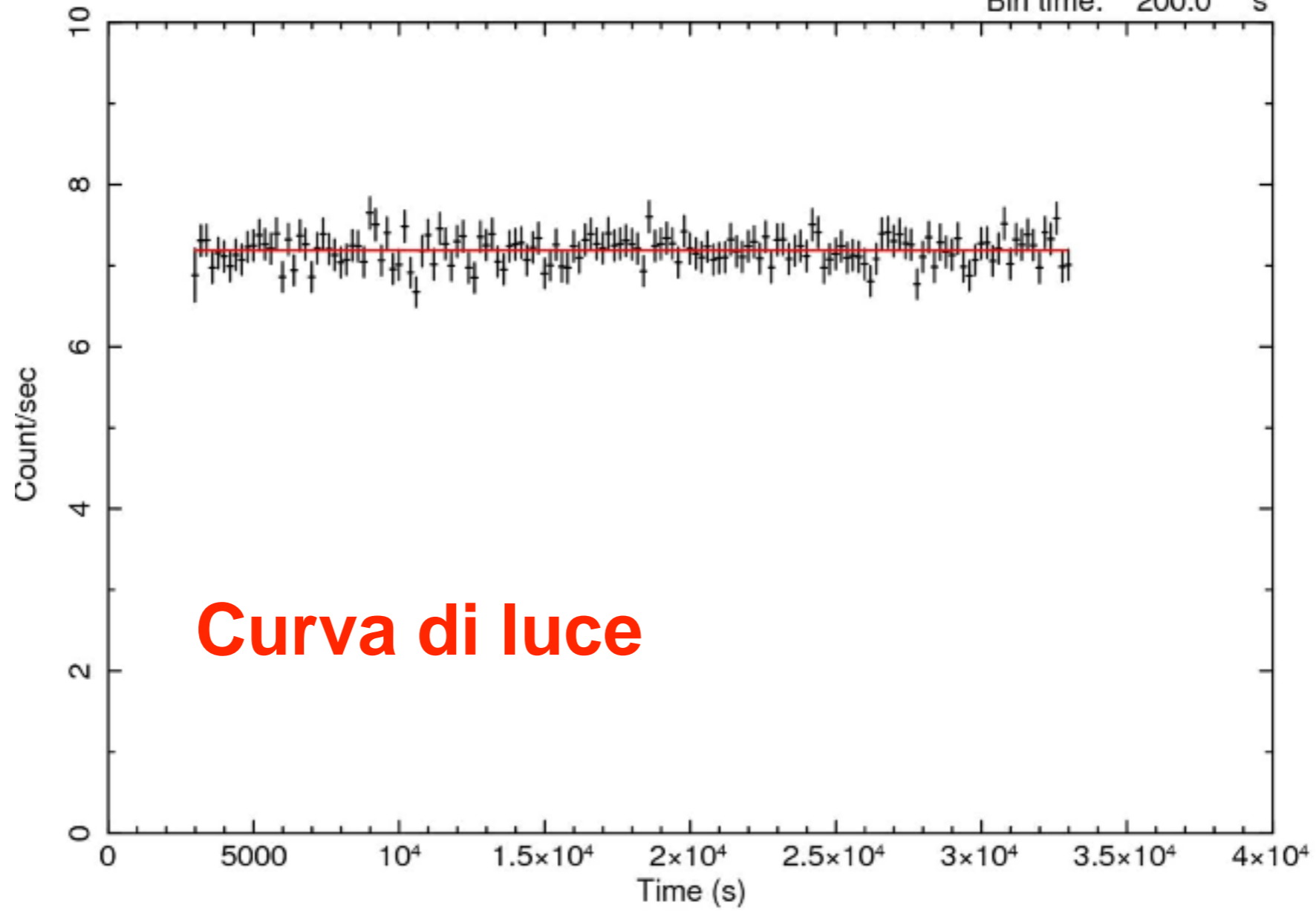


Imagine



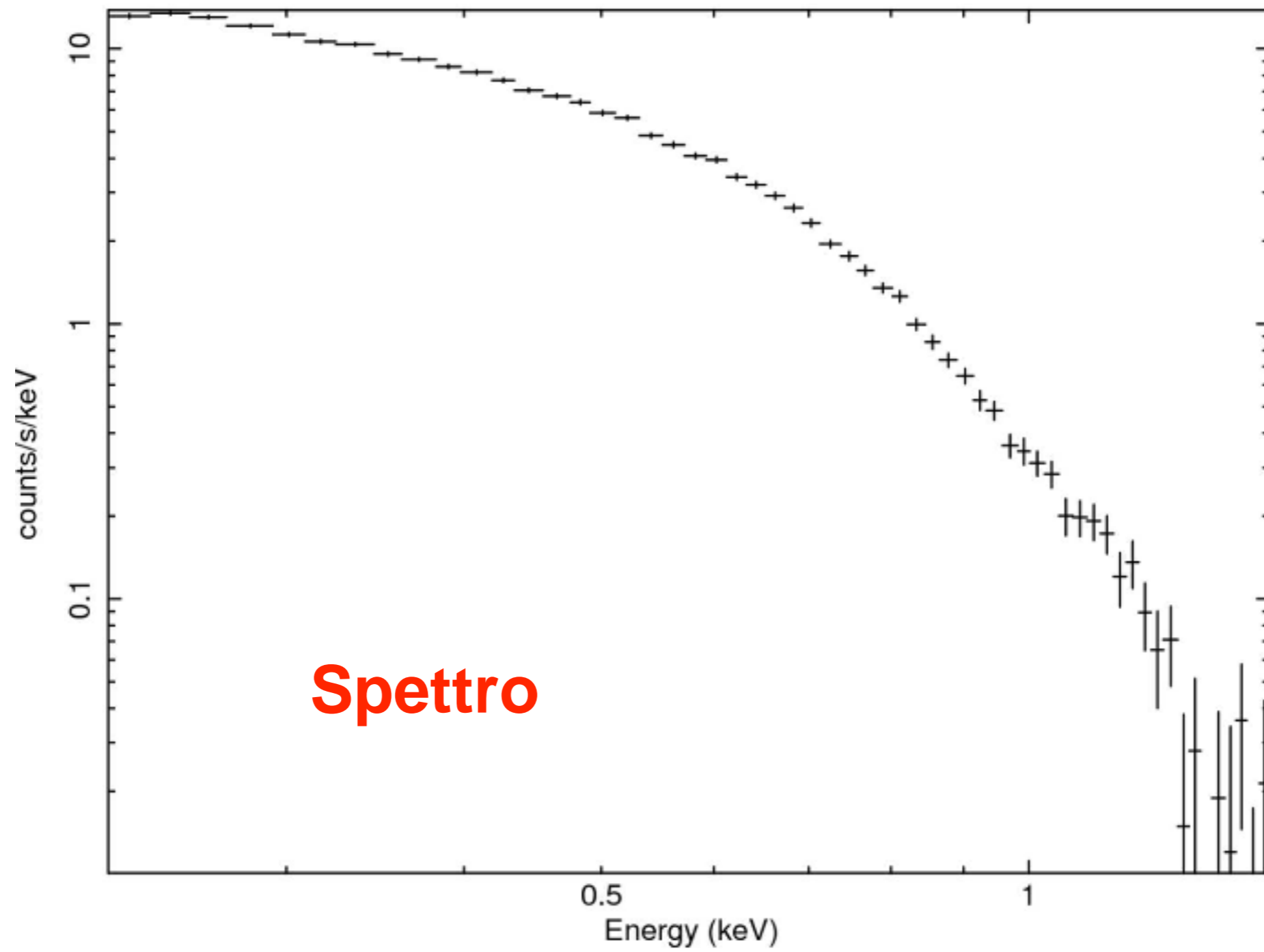
RXj 0720.4-3125

Bin time: 200.0 s



Curva di luce

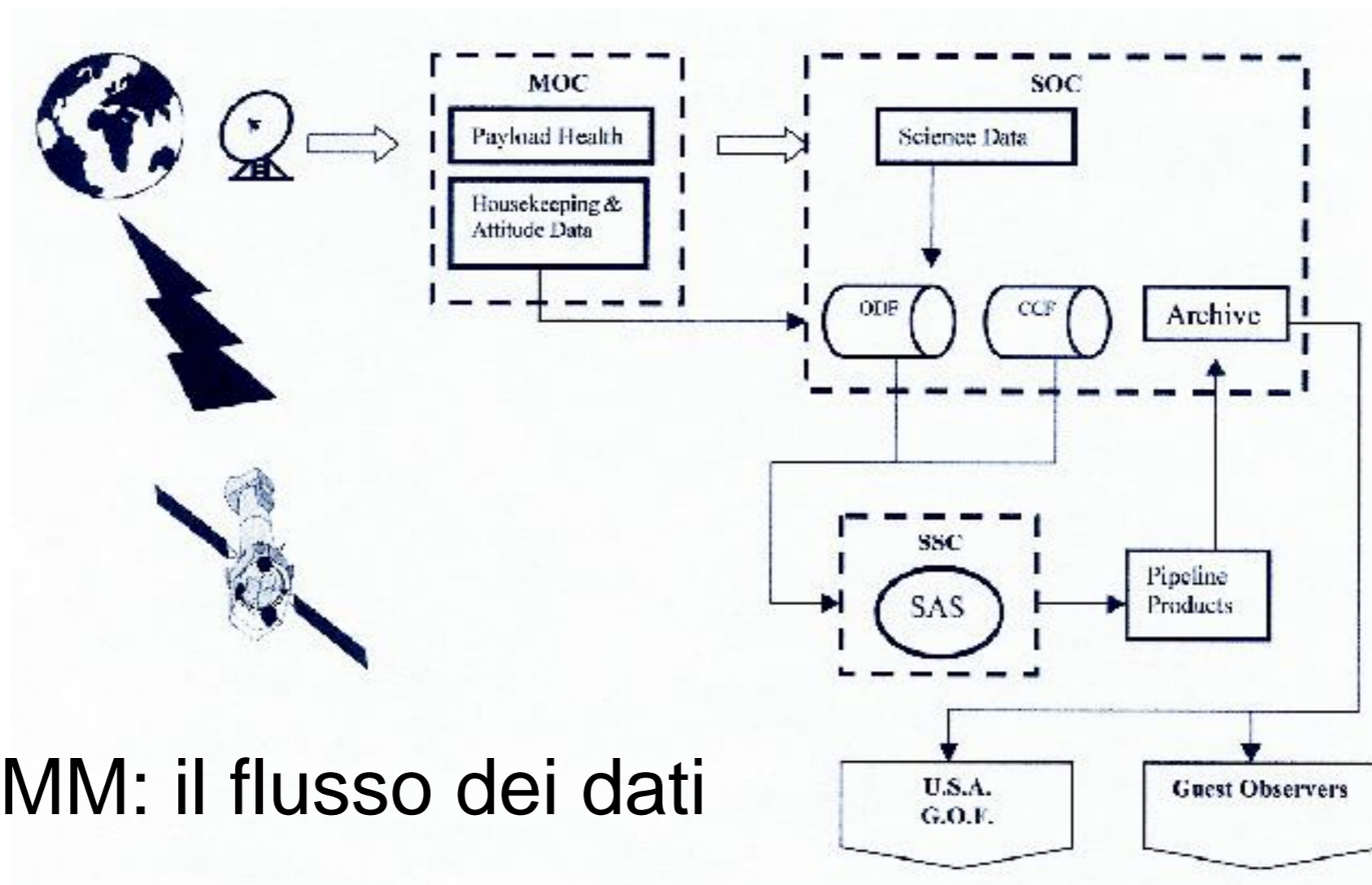
Start Time 12586 19:49:46:880 Stop Time 12587 4:09:46:880



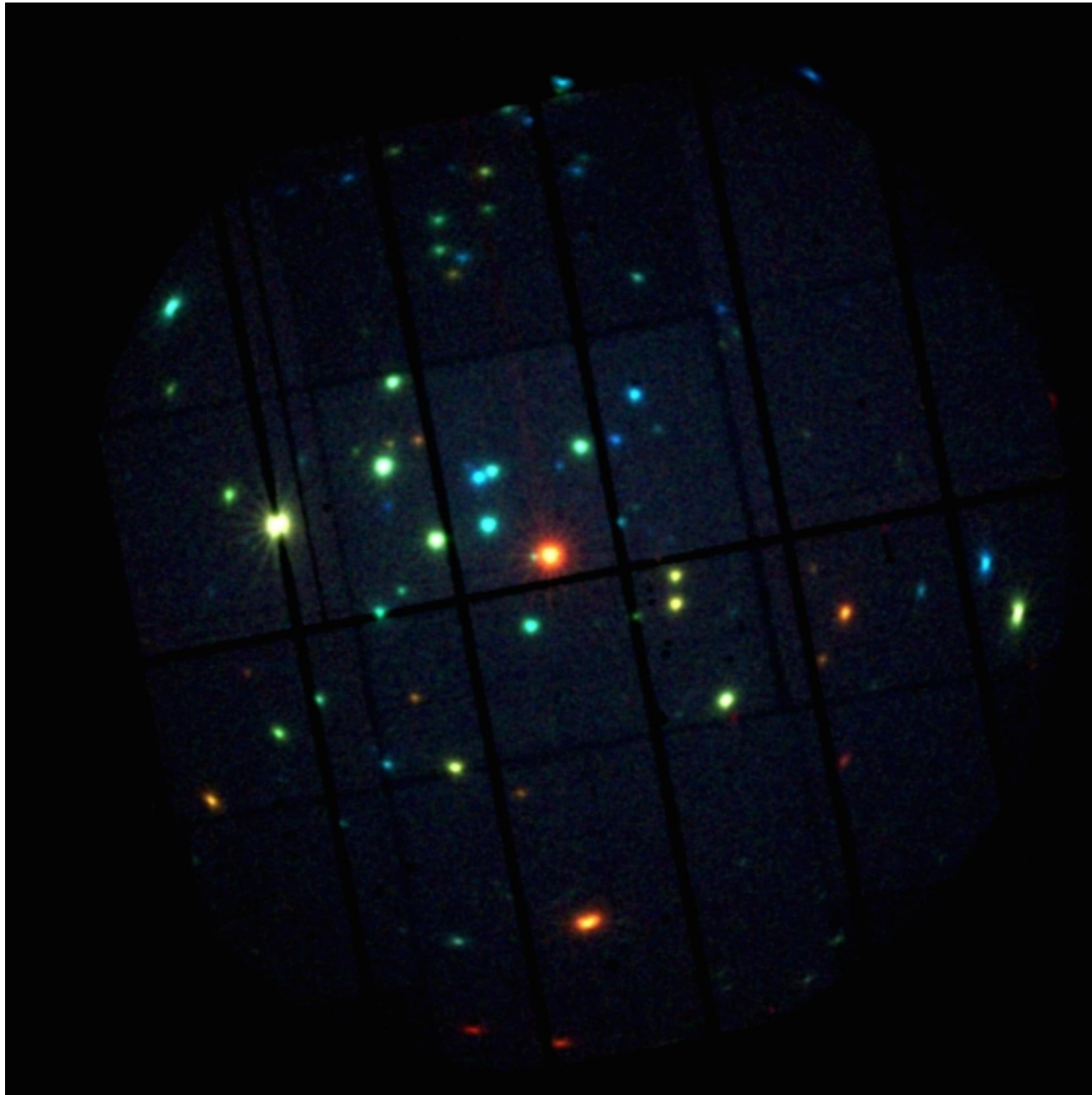
XMM e' un osservatorio spaziale

Bando per il tempo osservativo

Dati pubblici dopo 1 anno



La scienza *serendipita*



**Scoperte “fortuite”
in astronomia:
esempi famosi
(3 premi Nobel)**

L’astronomia X nasce
con 2 scoperte serendipite
nella stessa osservazione

The EPIC database: pointed observations

Pointed obs. → 3XMM catalogue*

794 deg² non-overlapping

531261 sources (373268 unique)

66728 with >1 detection (up to 44)

0.5-2 keV sensitivity $3(1) \cdot 10^{-15}$ cgs

2-10 keV sensitivity $1.5(0.8) \cdot 10^{-14}$ cgs

*by the XMM SSC

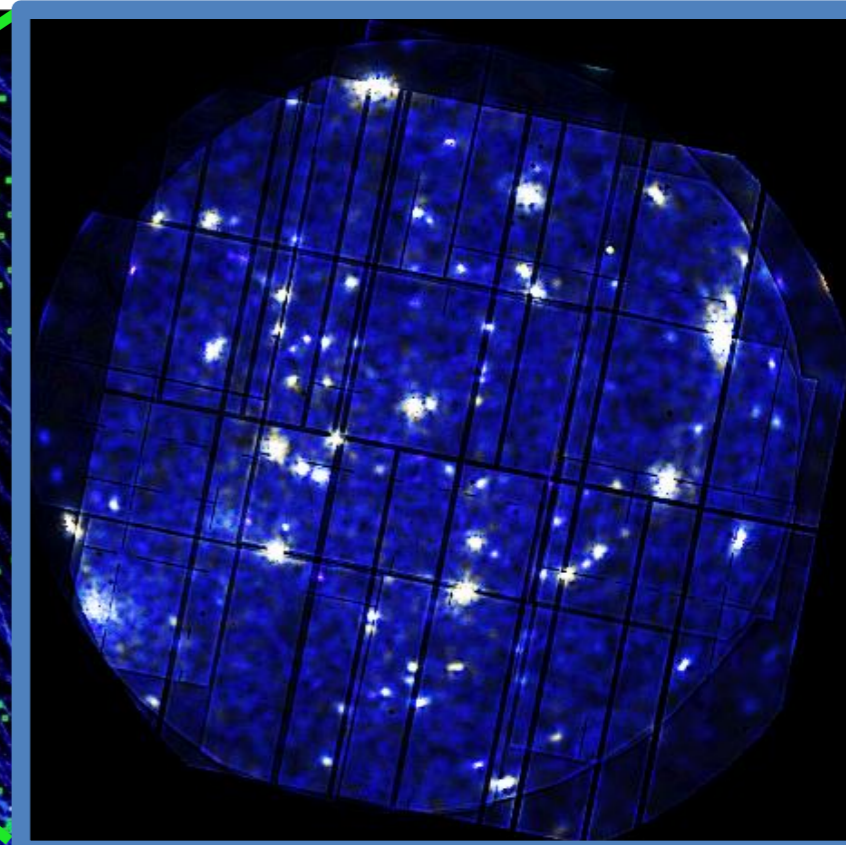


image by S.Rosen

The EPIC database: slew data

image by A. Read

Slew obs.

68% sky (22% >1 epoch)
7-10 s per epoch

Slew obs. → XSS catalogue

20163 sources (18400 unique)
950 with >1 detection (up to 8)
0.5-2 keV sensitivity $6 \cdot 10^{-13}$ cgs
2-10 keV sensitivity $2 \cdot 10^{-12}$ cgs

image by S.Rosen

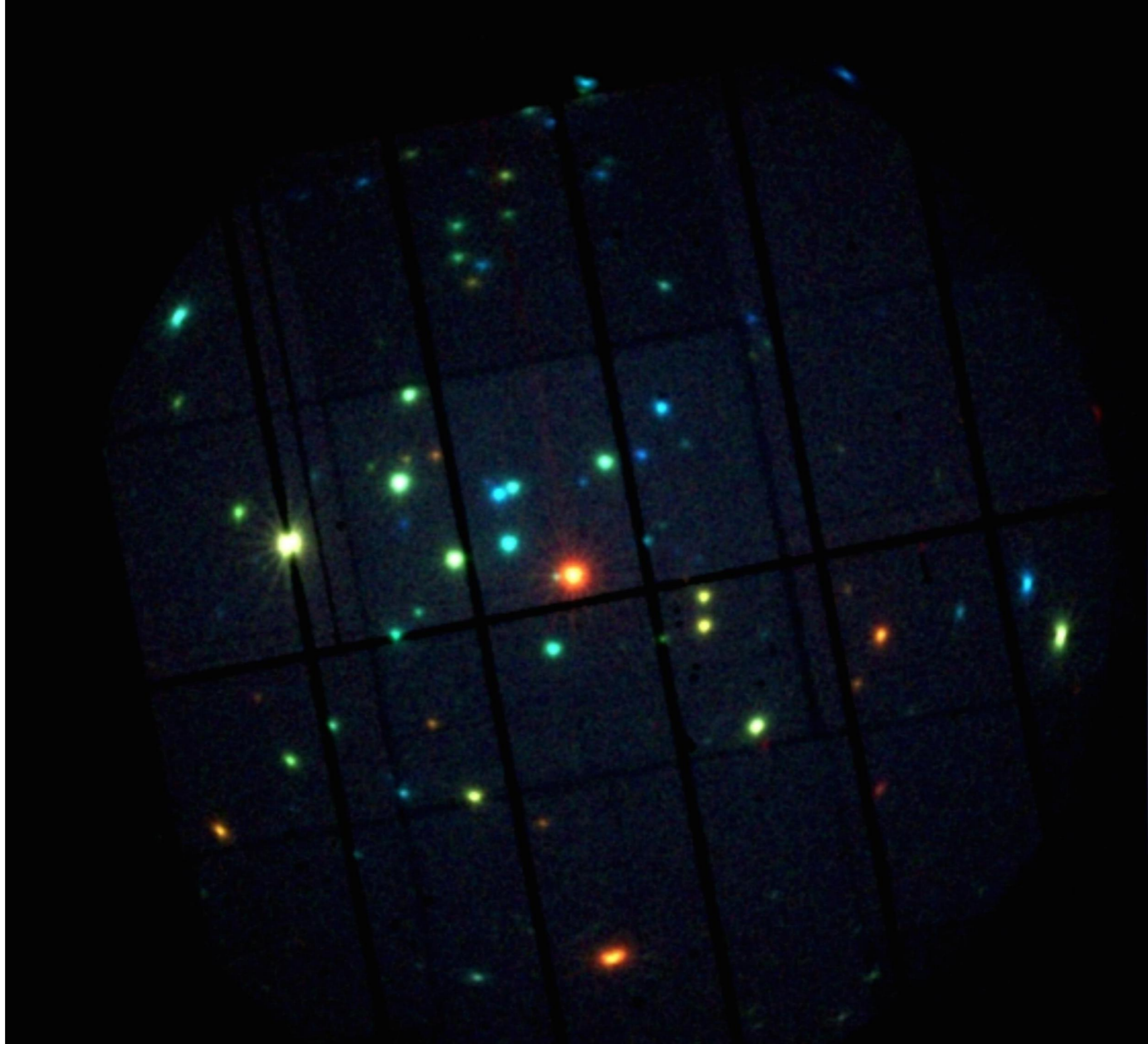


Resta qualcosa da fare?

EXTraS

Exploring the **X**-ray **T**ransient and variable **S**ky

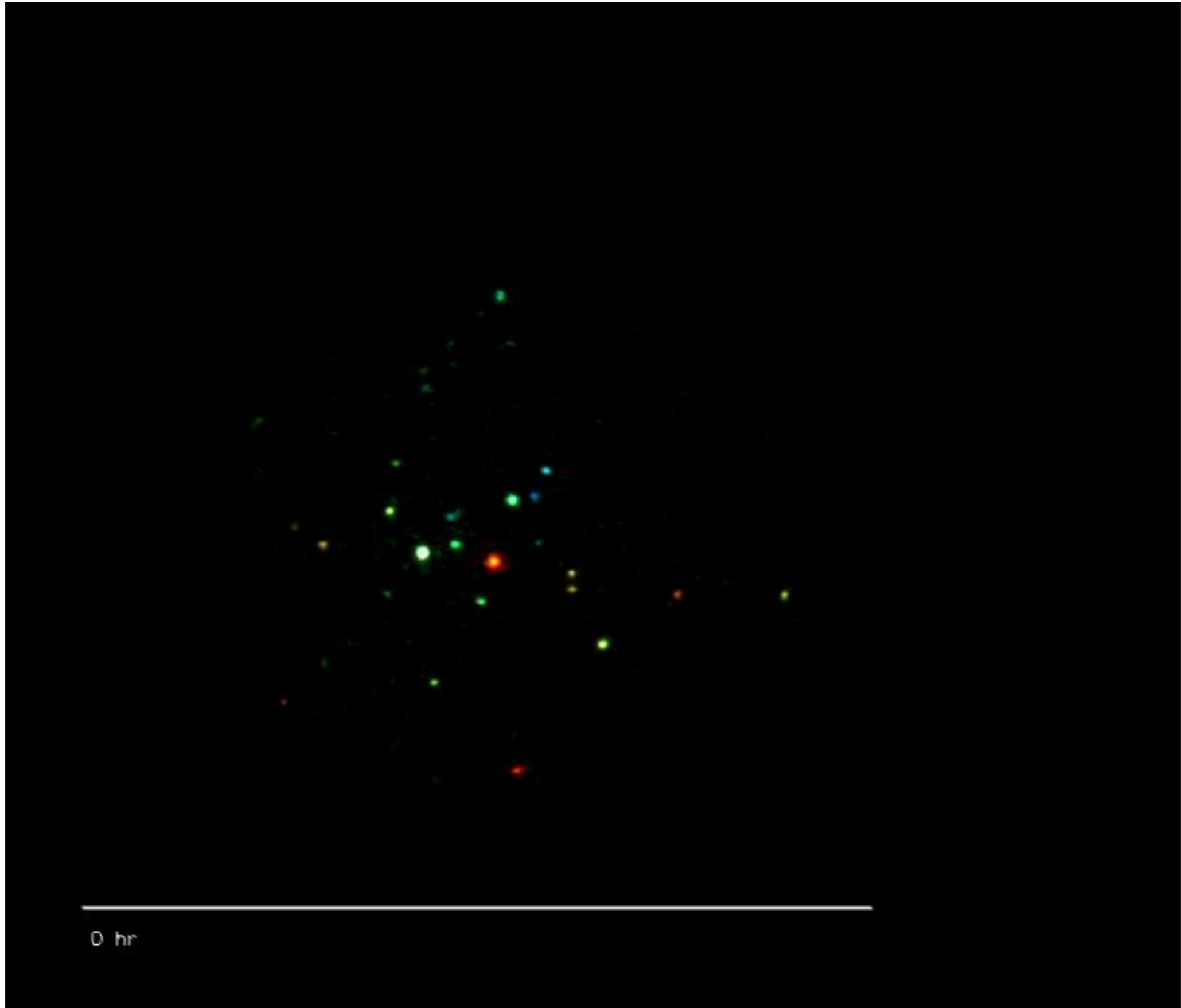




Cosa e' stato "catalogato" di queste sorgenti?

Video time: droxo movie

http://xmm.esac.esa.int/external/xmm_science/gallery/images/droxomovie.mpg



0 hr

Astronomia nel dominio temporale

“Variability pervades the cosmos.

Studies of variability dominate research in astronomy and astrophysics and are so common that very many groups, projects and instruments are dedicated to the examination of just one form of variability, or one aspect of its diverse manifestations.”

manifest of IAU Symposium 285, "New Horizons in Time Domain Astronomy",

University of Oxford September 19-23, 2011

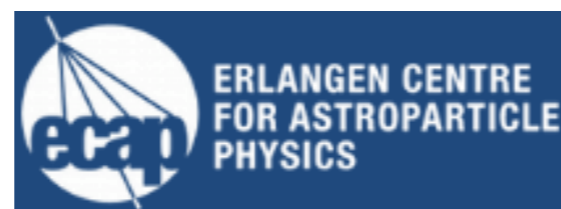
Come e' nato EXTraS

- variabilita' cruciale per capire sorgenti
- variabilita' presente in tutte le sorgenti in banda X
- dati serendipiti EPIC: miniera ancora da sfruttare

L'occasione

il bando "Spazio" di FP7

Il consorzio EXTraS



Brevissima storia di EXTraS

- idea: estate 2012
- costituzione consorzio: settembre 2012
- proposta sottomessa il 26 novembre 2012
- progetto selezionato a marzo 2013
- contratto firmato agosto 2013
- inizio lavori gennaio 2014

EXTraS in breve

The most sensitive and complete investigation of variability in the soft X-ray sky

explore the **serendipitous** content of the **XMM-Newton/EPIC** database in the **time domain**

make it **available and easy to use** to the whole **community**.

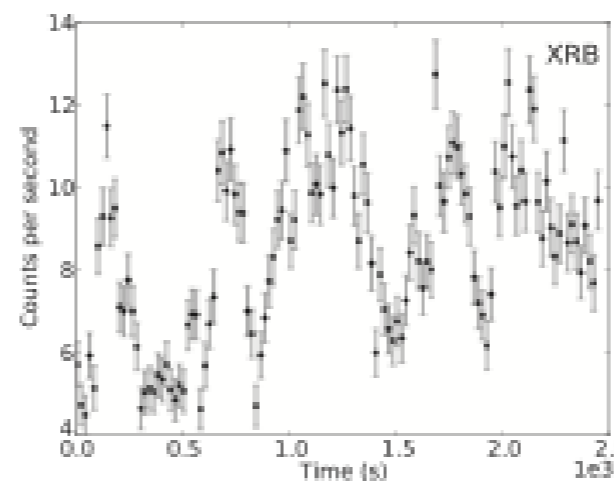
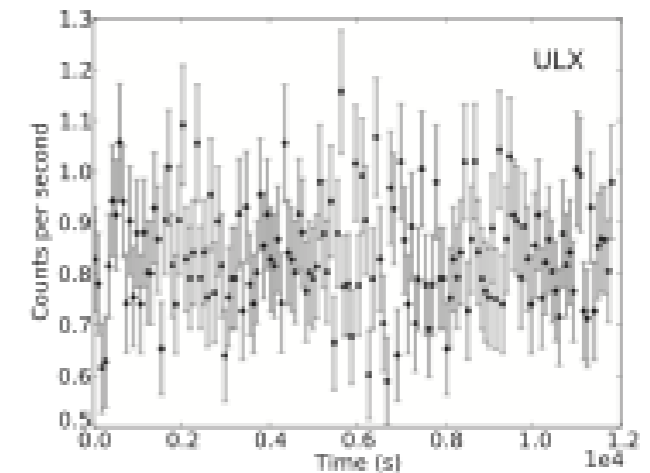
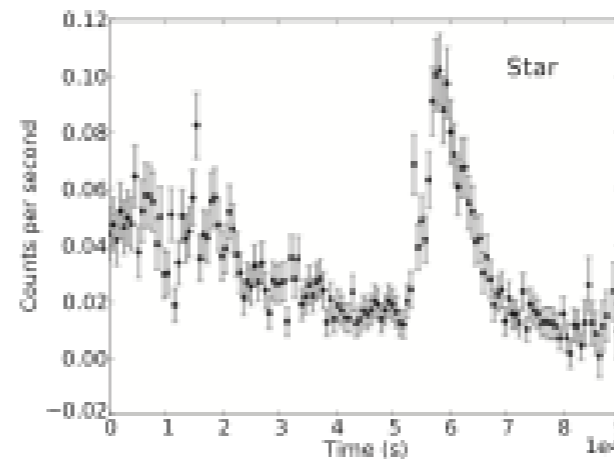
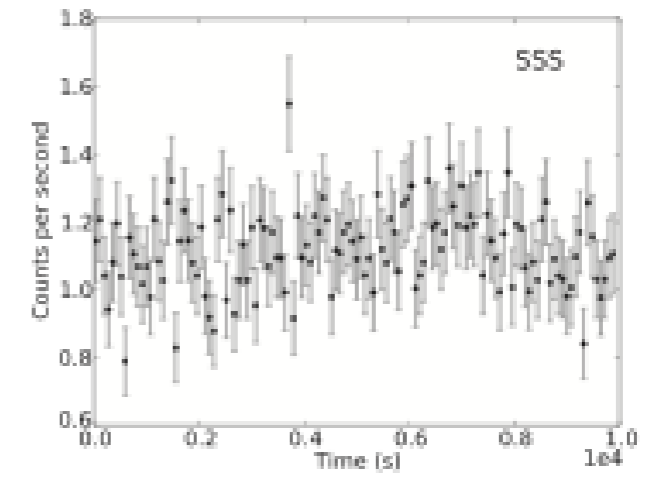
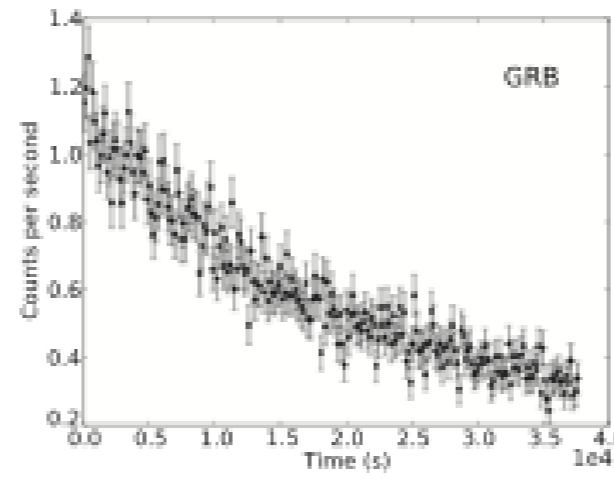
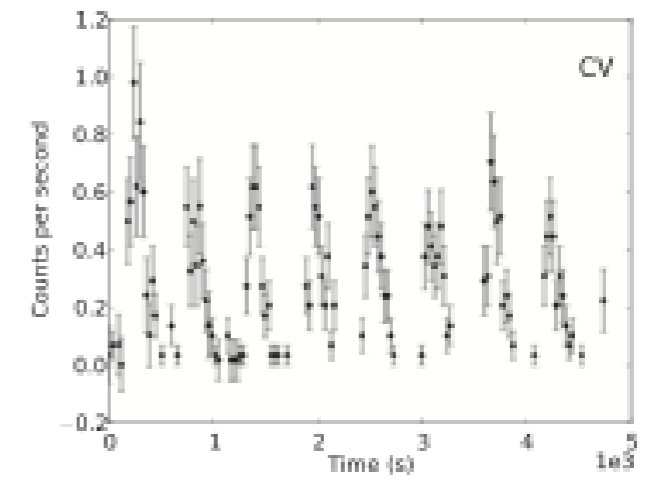
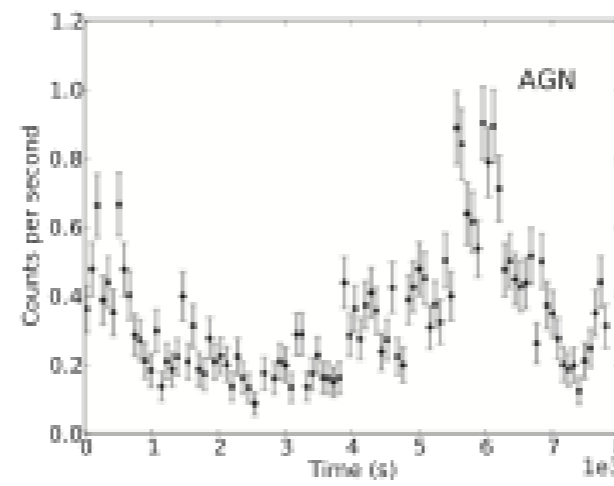
allow to do **more, new science**

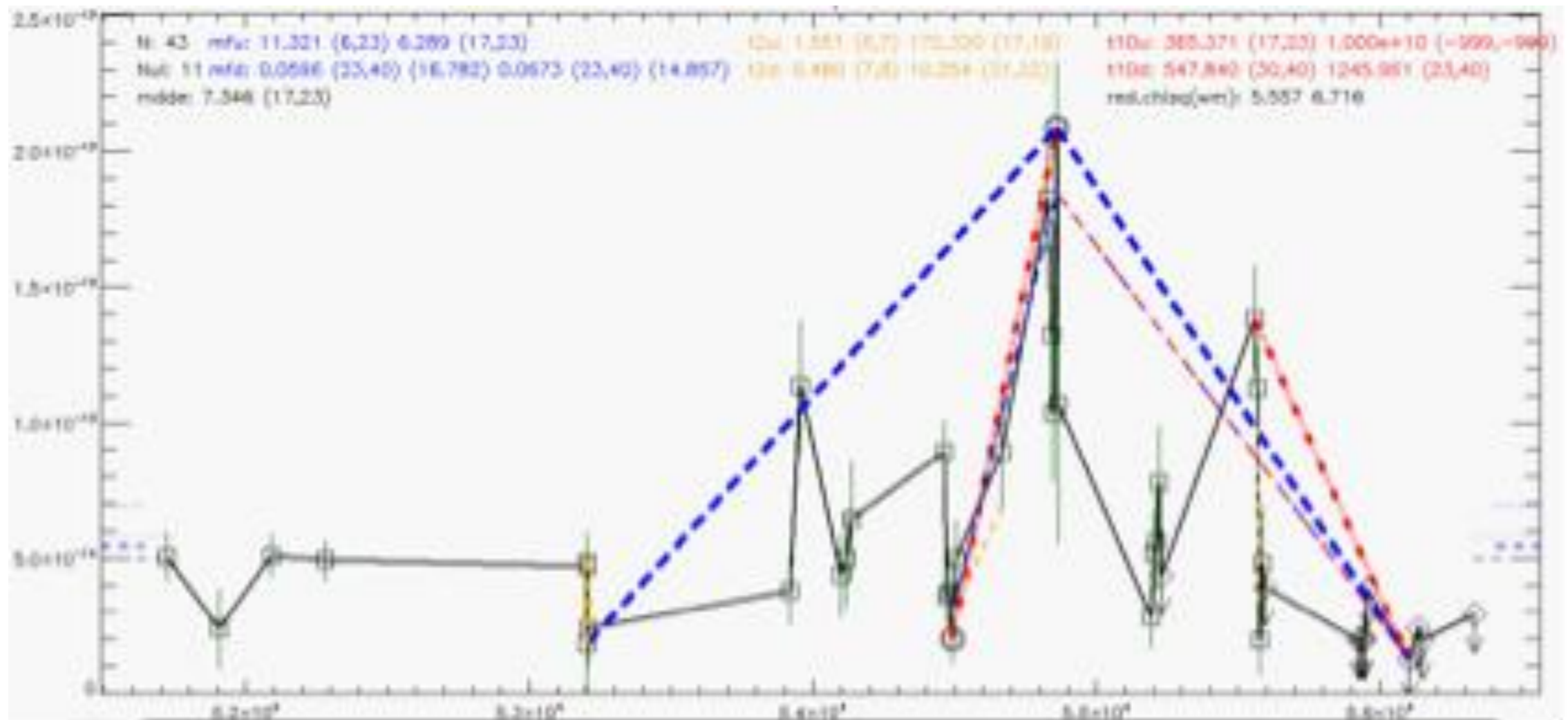
Cosa faremo in EXTraS

- **Variabilita' non periodica**
- **Variabilita' periodica (pulsazioni)**
 - **Transienti**
- **Variabilita' su lungo tempo scala**

Diversi esempi
di curve di luce
con diverse variabilita'

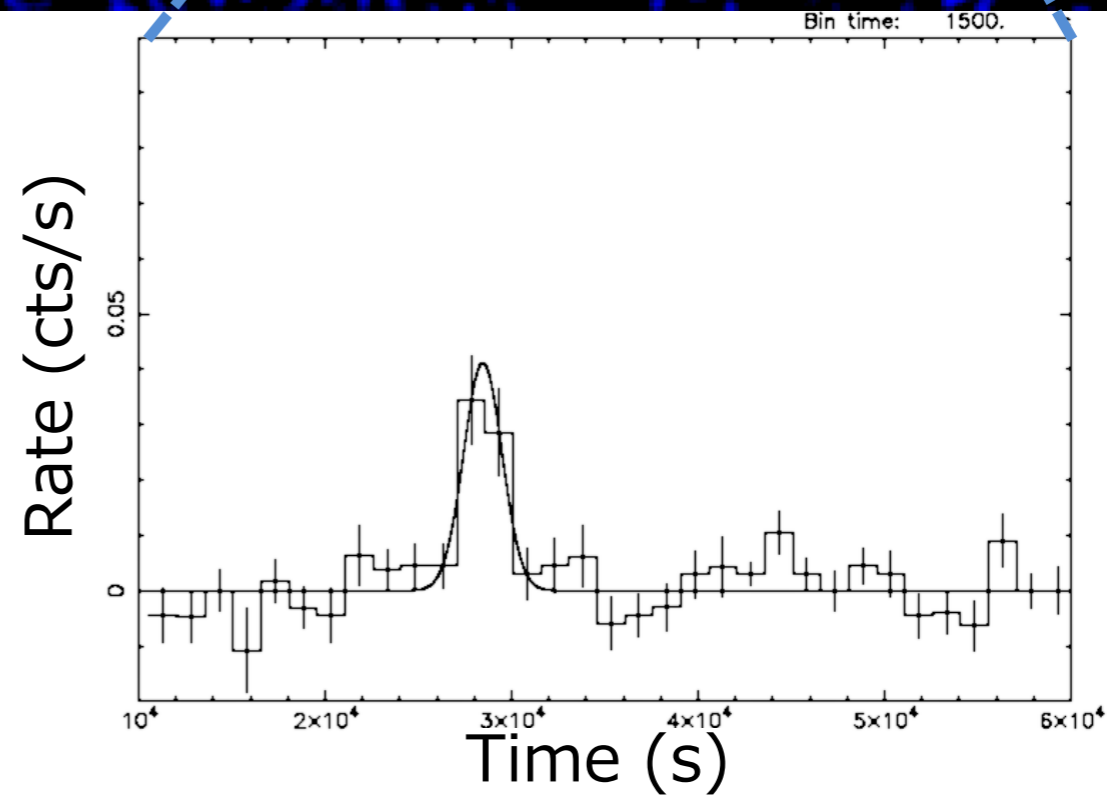
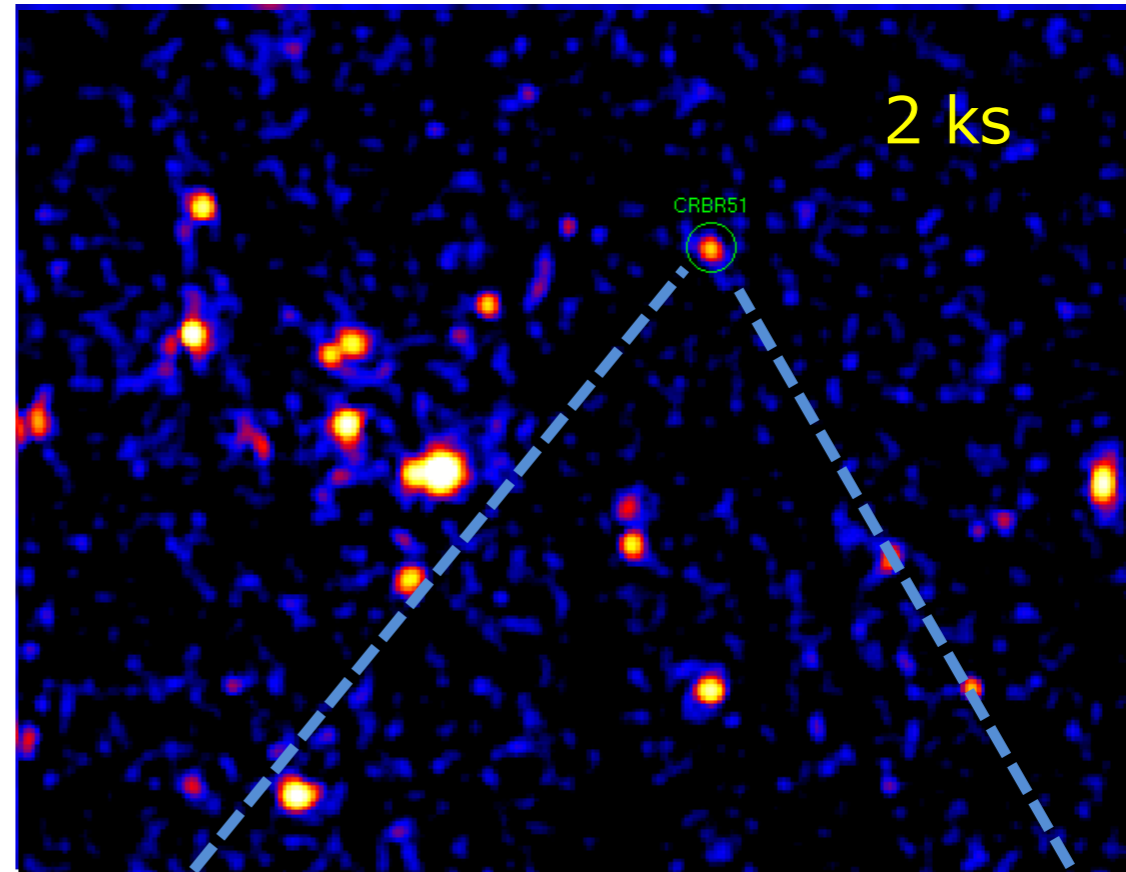
tempi scala
da minuti a ore



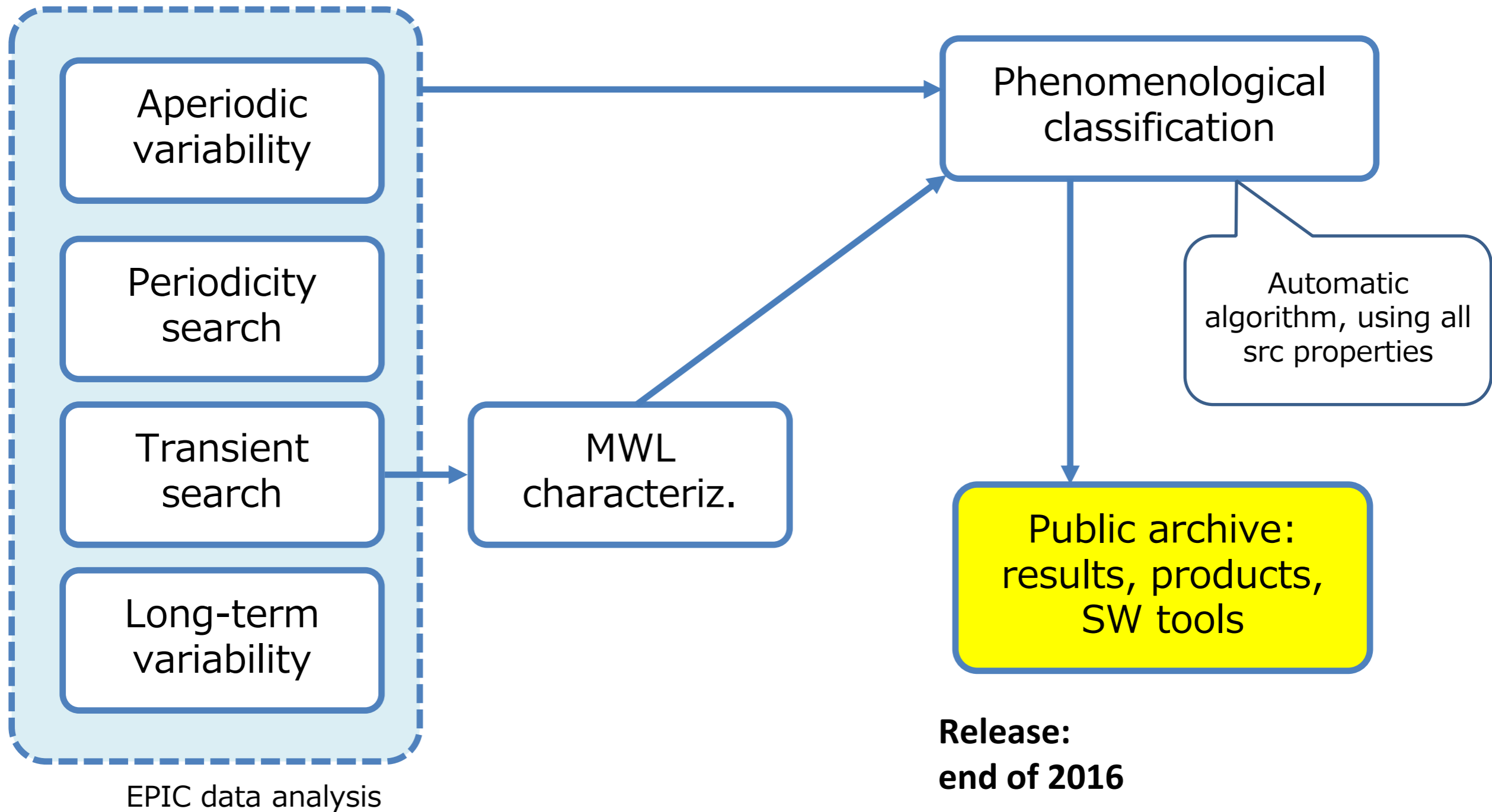


Una curva di luce estesa su 15 anni

Una sorgente “transiente”
visibile per soli 30 min

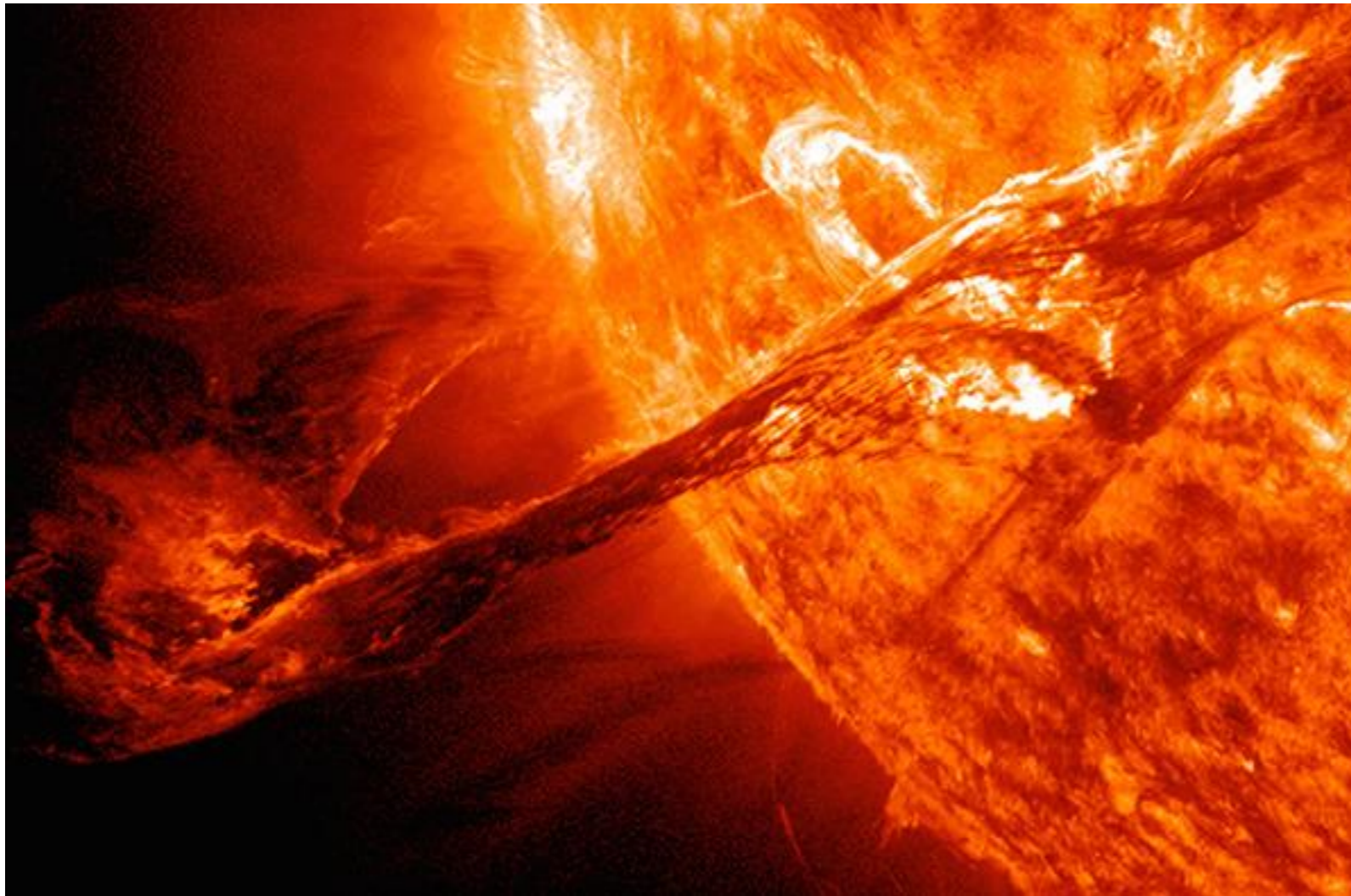


EXTraS: the output



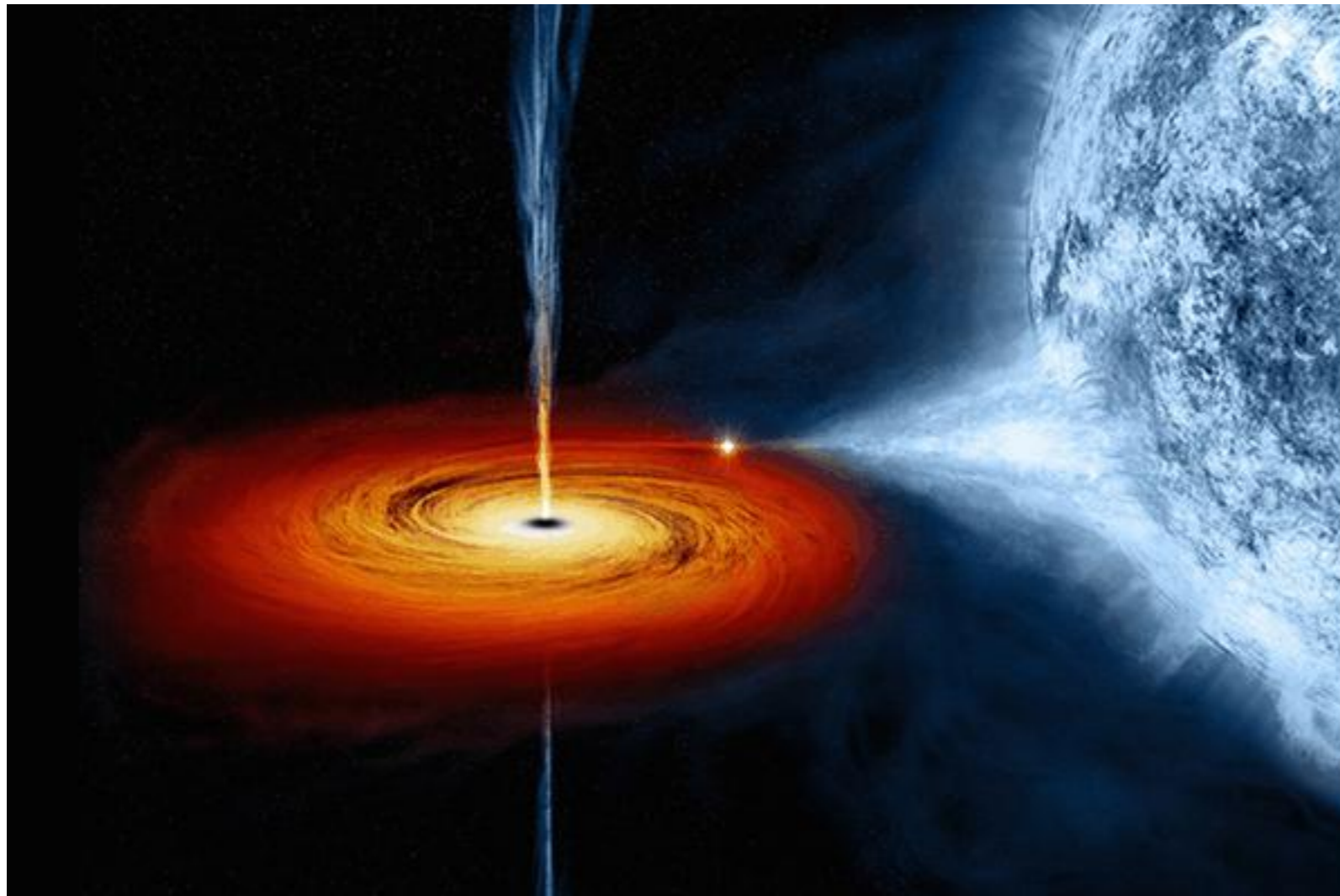
Le sorgenti di EXTraS

Tutte le sorgenti in banda X sono variabili



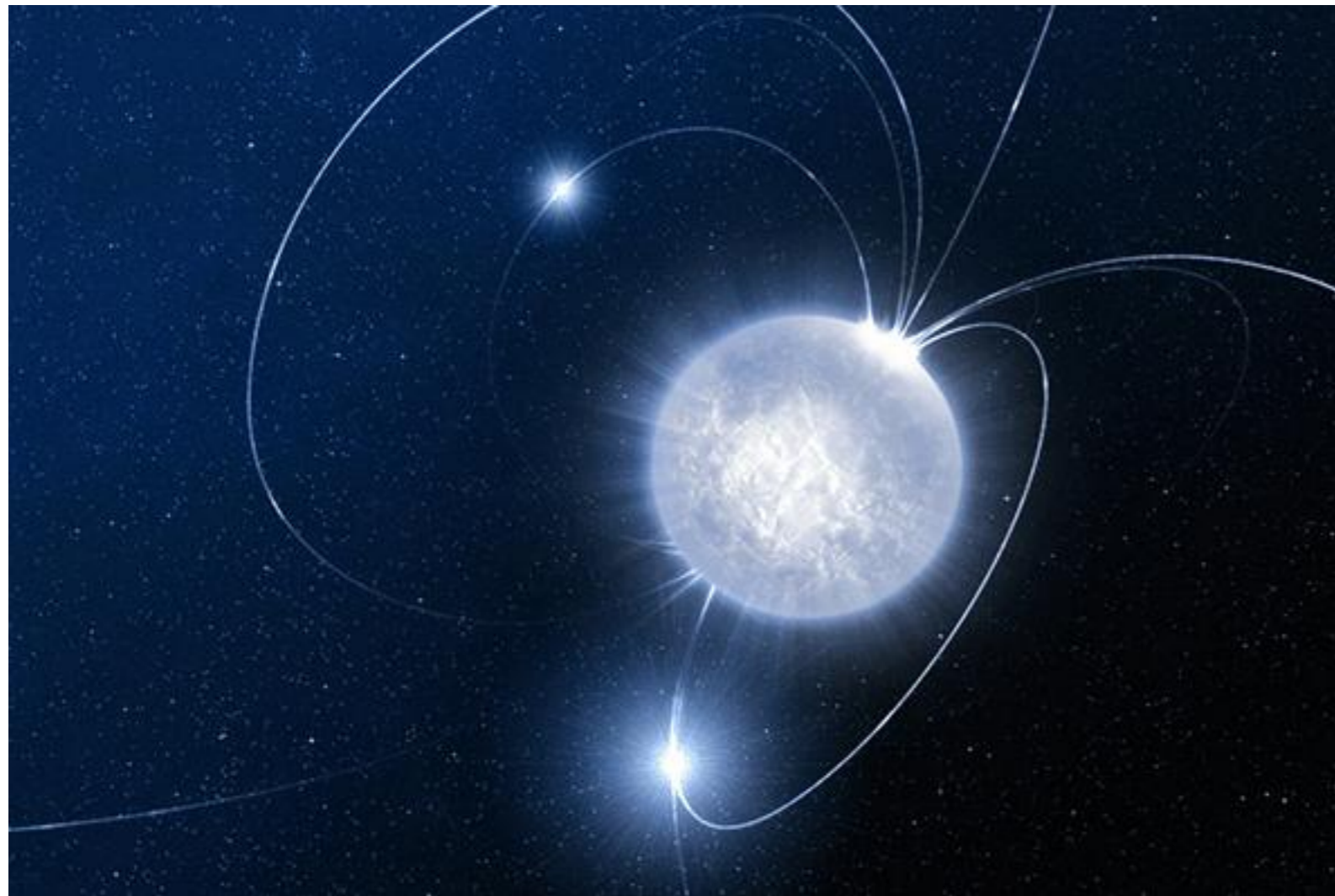
Stellar flares

X-ray flares from magnetically active, late-type stars, either isolated or in binary systems.



Transient X-ray binaries

Black holes, neutron stars or white dwarfs accreting matter from their stellar companion.

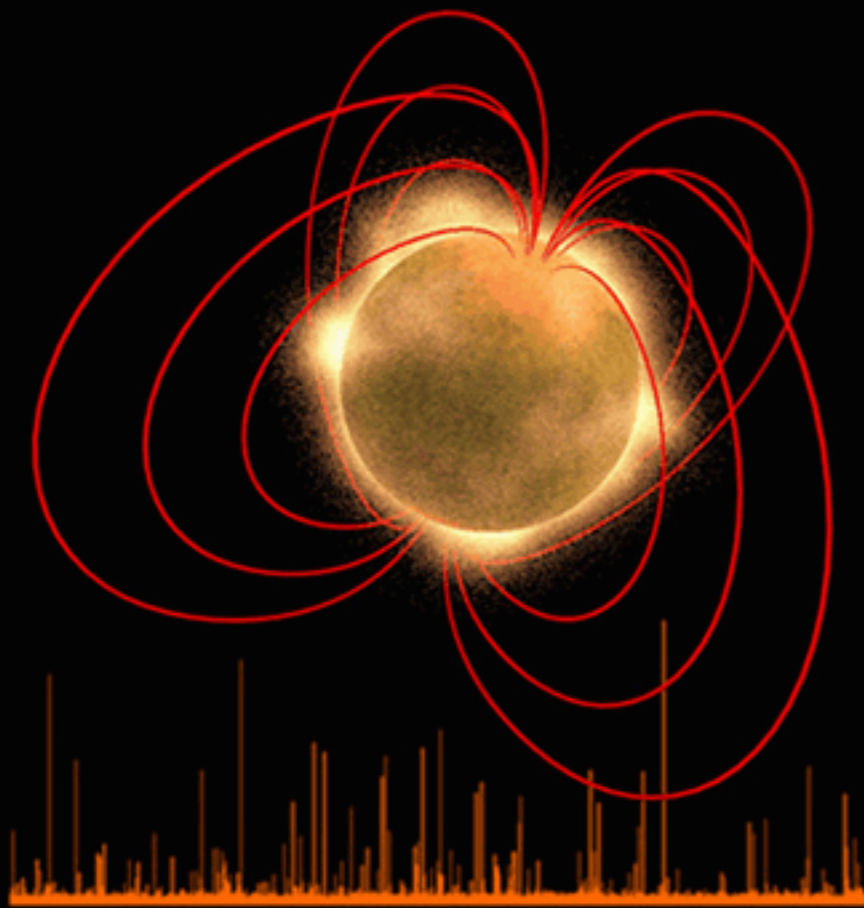


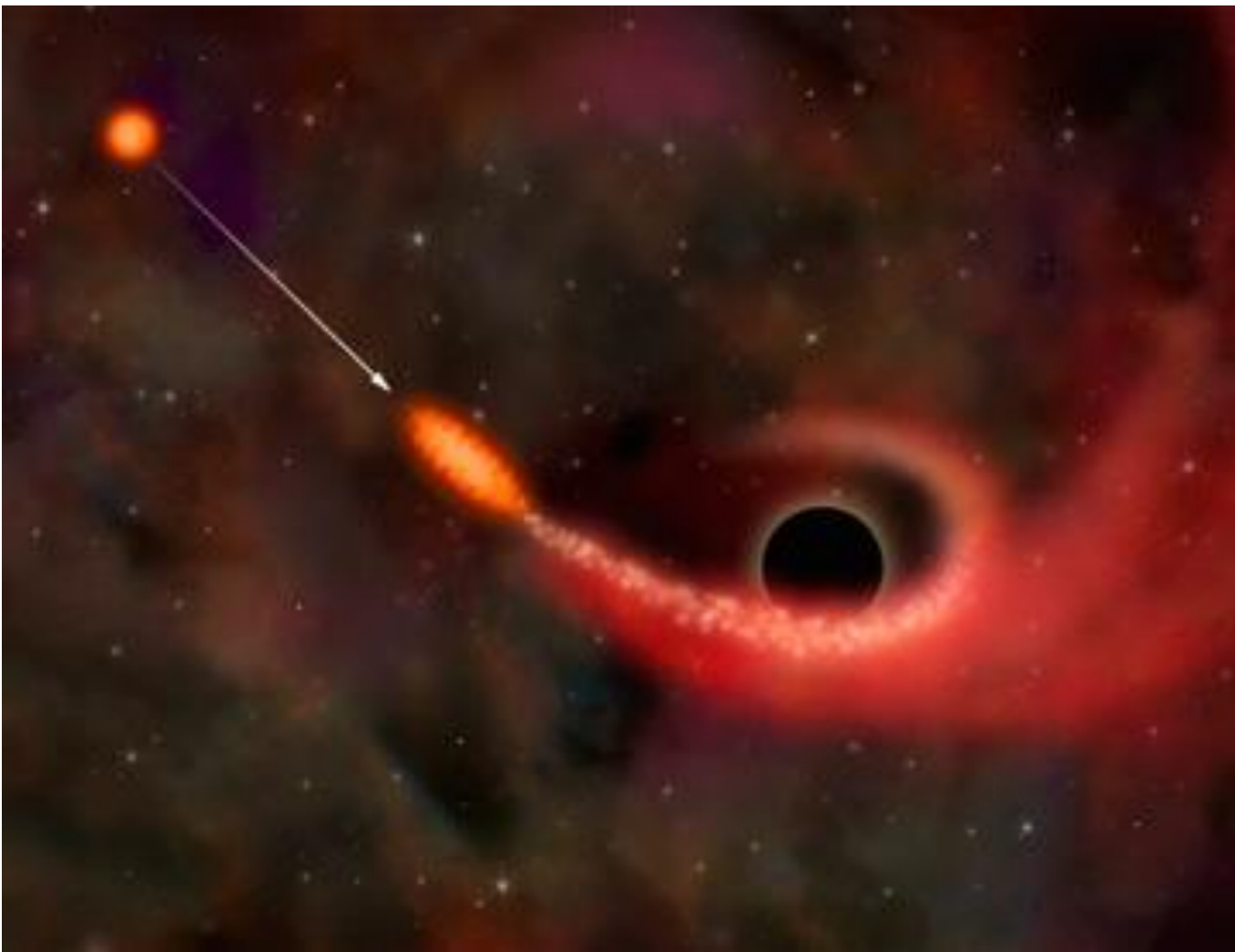
Magnetar

Magnetar is a type of neutron star with an extremely powerful magnetic field.

Soft gamma-ray repeaters (SGRs)

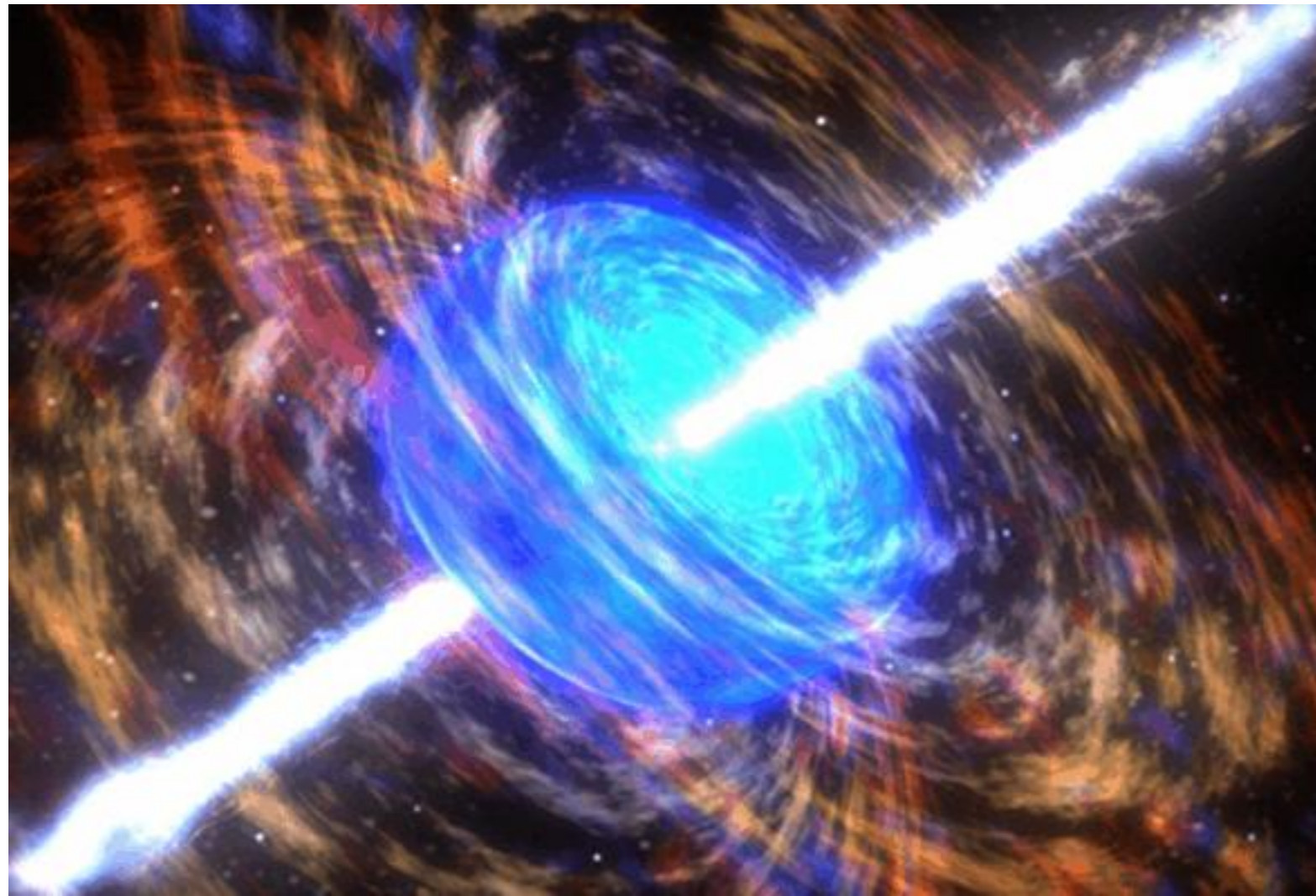
X-ray sources believed to be powered by magnetars, i.e. neutron stars with the strongest magnetic field in the Universe.





Tidal Disruption Event

Disruption of a star
by a massive black hole



Gamma-ray bursts (GRBs)

The most powerful cosmic explosions, likely produced by the collapse of massive stars to black holes or by the coalescence of two neutron stars.



Supernova X-ray flashes

Produced by the supernova shock emerging from the exploding star.



Blazar flares

Gamma-ray flares produced by the jets of supermassive black holes at the centre of galaxies.

Some XMM-Newton history



- Mission was built for 2.25 years
- Mission was designed to be compatible with 10 year lifetime
- Before launch there was some 100 kg mass margin that was used to 'top up' the fuel
- Mission was built to be 24/7 controlled from ground – no intelligence – no significant command stack
- We launched with only 3 quadrants of EPIC-PN working ! (There's nothing like a good shake)

- I started working on XMM-Newton in 1988 – 26 years ago →
Knowledge management is an issue
 - Usually KM documents a design and what the system CAN do – not what it COULD be made to do