

## LESSON 1 - The ultraviolet catastrophe and Planck's formula

### Part 1 – Revision: e.m. waves

Match the following terms with their definition: *(do this part on Spark, so you can check immediately your answers).*

1. Electromagnetic wave	A. The range of wavelengths or frequencies over which electromagnetic radiation extends.
2. Wavelength	B. The total number of vibrations or oscillations per unit time.
3. Frequency	C. The power per unit area delivered by a wave.
4. Electromagnetic spectrum	D. The distance between successive crests of a wave.
5. Intensity (of a wave)	E. A wave propagated by the periodic variation of intensities of electric and magnetic fields.

### Part 2 – Video watching: ultraviolet catastrophe and Planck's formula

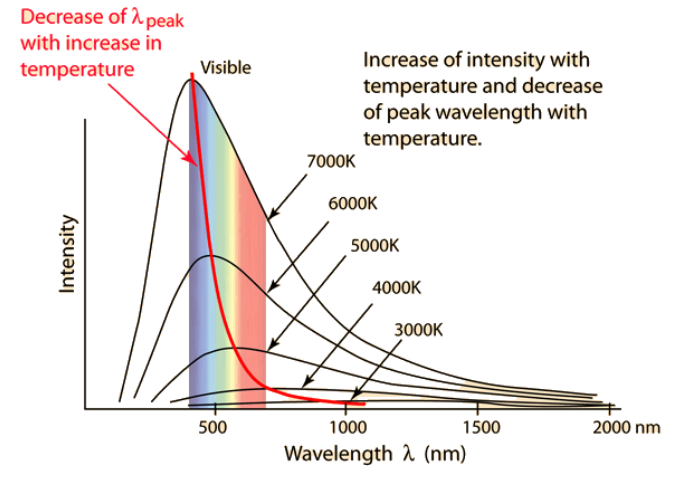
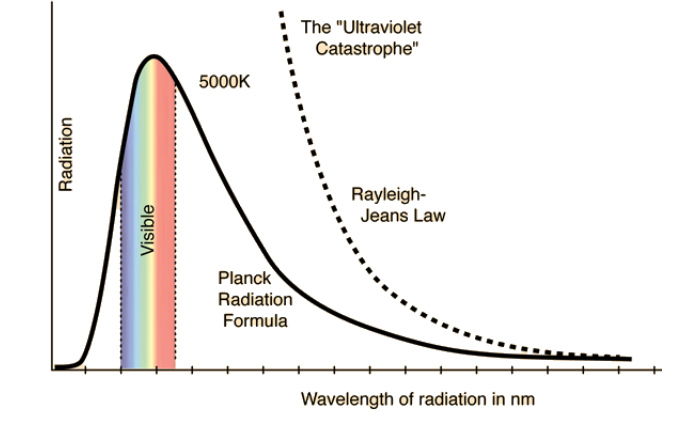
Watch the video “Quantization of Energy Part 1: Blackbody Radiation and the Ultraviolet Catastrophe” and complete the following sentences. *The video is also on Spark; we will check your answers in our next video lesson.*

1. A blackbody emits e.m.radiations of all \_\_\_\_\_
2. The light \_\_\_\_\_ by the sun matches the \_\_\_\_\_ for 5778K. Most of the light that we receive from the sun is in the \_\_\_\_\_ spectrum.
3. The blackbody spectrum depends only on \_\_\_\_\_
4. The wavelength that is emitted with \_\_\_\_\_ intensity shifts left as temperature increases. This maximum will move into the visible spectrum at around \_\_\_\_\_ K and above. This is why very hot objects appear to \_\_\_\_\_, like a hot \_\_\_\_\_, light \_\_\_\_\_ filament or the \_\_\_\_\_ and other stars.
5. According to classical theory, shorter wavelengths were not predicted to \_\_\_\_\_ in intensity.
6. Classical models don't match \_\_\_\_\_, otherwise every time you use the oven you would get blasted with \_\_\_\_\_
7. Max \_\_\_\_\_ solved the problem introducing a concept called \_\_\_\_\_
8. Planck proposed that the \_\_\_\_\_ of the atoms must be quantized, meaning that they can only possess specific \_\_\_\_\_ values.
9. Planck's expression for blackbody radiation is  $E = nhf$ , where  $n$  can be any \_\_\_\_\_,  $h$  is called the Planck's \_\_\_\_\_ ( $h = 6.626 \cdot 10^{-34} \text{ J} \cdot \text{s}$ ) and  $f$  is the \_\_\_\_\_ of radiation.

10. Energy appears to be \_\_\_\_\_ to macroscopic beings, such as humans, but it is quantized on the \_\_\_\_\_ of the scale (not observable).

**Part 3 – Describe pictures and formulas**

In the following table, you can find pictures and formulas related to two important laws. They are not explicitly named in the video, but professor Dave talks about them anyway. Try to describe them: *do this on Spark, so I can eventually check if your answers are right. We will correct them anyway during our next video lesson.*

	<p><b>Wien's displacement law:</b></p> $\lambda_{peak} \cdot T = constant$
	<p><b>Rayleigh-Jeans law:</b></p> $I \propto \frac{T}{\lambda^4}$

**Part 4 – Key points**

Watch again the video and rearrange in the right order the key points of the lesson: *do this part on Spark, so you can immediately check your answers.*

- Planck's quantization of energy
- Definition of a blackbody
- Ultraviolet catastrophe and Rayleigh-Jeans formula: classical Physics is not able to explain blackbody radiation
- Description of blackbody radiation and Wien's displacement law